

*ДВОЙНОЙ БЕТА-РАСПАД: современное  
состояние и перспективы*

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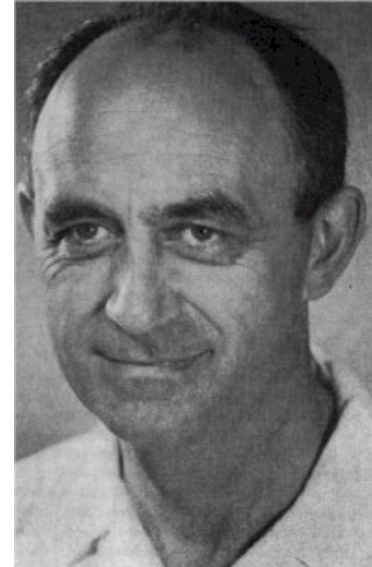
# ПЛАН

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- Историческое введение
- Современное состояние (**NEMO-3, CUORICINO**)
- Эксперименты следующего поколения
- Заключение

# I. Historical introduction

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Neutrino was introduced by **W. Pauli** in 1930

$\beta$ -decay theory (weak interaction) was formulated by **E. Fermi** in 1933:

$$(A, Z) \rightarrow (A, Z+1) + e^- + \bar{\nu}$$

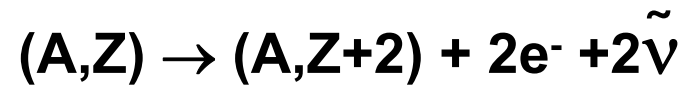
$$(A, Z) \rightarrow (A, Z-1) + e^+ + \nu$$

# The birth of double beta decay

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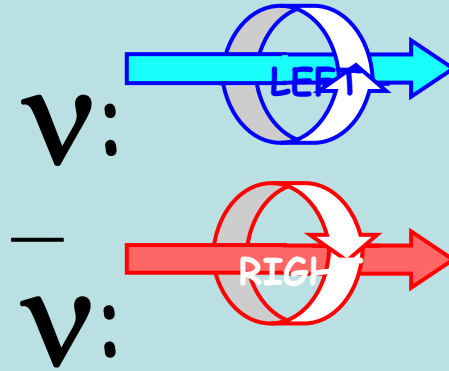


- $2\beta(2\nu)$  decay was introduced by **M. Goeppert-Mayer** in 1935:



$$(T_{1/2} \sim 10^{21}-10^{22} \text{ y})$$





$$\nu \neq \bar{\nu}$$

$$\nu = \bar{\nu}$$

Dirac



Majorana

=>1937



Dirac particle



Majorana particle



# The birth of neutrinoless double beta decay

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- $2\beta(0\nu)$  decay was introduced by W.H. Furry in 1939:



$$(T_{1/2} \sim 10^{15}-10^{16} \text{ y})$$

[Parity violation was not known at that time!]

# First experiments

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- **1948** – first counter experiment (Geiger counters,  $^{124}\text{Sn}$ ;  $T_{1/2}(0\nu) > 3 \cdot 10^{15} \text{ y}$ )
- **1950** – **first evidence** for  $2\beta 2\nu$  decay of  $^{130}\text{Te}$  in first geochemical experiment:  
 $T_{1/2} \approx 1.4 \cdot 10^{21} \text{ y}!!!$
- **1950-1965** – a few tens experiments with sensitivity  $\sim 10^{16}-10^{19} \text{ y}$
- **1966-1975** – in 3 experiments sensitivity to  $0\nu$  decay reached  $\sim 10^{21} \text{ y}!!!$

# 1957 – situation is changed!

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- **P** and **C** violation
- **V-A** structure of weak interaction
- **Helicity** of  $\nu(\bar{\nu})$  is  $\sim 100\%$



**$2\beta(0\nu)$ -decay is suppressed (if even possible?)**

and  **$T_{1/2}(0\nu) > T_{1/2}(2\nu)$**



# Best results in 1966-1975

- $T_{1/2}(0\nu; ^{76}\text{Ge}) > 5 \cdot 10^{21} \text{ y}$ ; Ge(Li) detector, 1973 (E. Fiorini et al.)



- $T_{1/2}(0\nu; ^{48}\text{Ca}) > 2 \cdot 10^{21} \text{ y}$ ; streamer chamber + magnetic field + plastic scint., 1970 (C. Wu et al.)
- $T_{1/2}(0\nu; ^{82}\text{Se}) > 3.1 \cdot 10^{21} \text{ y}$ ; streamer chamber + magnetic field + plastic scint., 1975 (C. Wu et al.)

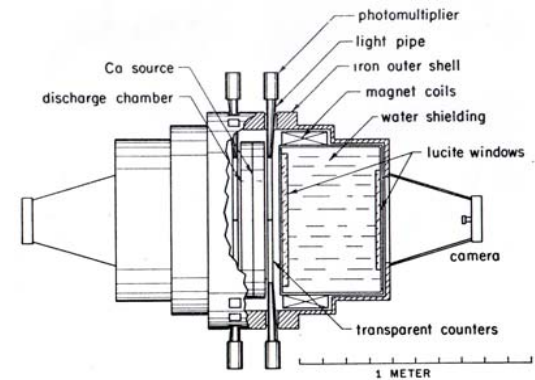


Fig. 3. Cutaway drawing of double beta decay apparatus.

- Geochemical experiments with  $^{130}\text{Te}$ ,  $^{128}\text{Te}$ ,  $^{82}\text{Se}$  ( $2\nu$  measurements:  $\sim 10^{21}$ ,  $\sim 10^{24}$  and  $\sim 10^{20}$  y)

## Main achievements in 1976-1987

- $2\beta 2\nu$  decay was first time detected in direct (counting) experiment  $\Rightarrow$

$$T(^{82}\text{Se})_{1/2} = 1.1^{+0.8}_{-0.3} \cdot 10^{20} \text{ y}$$

(35 events; TPC, **1987**,  
S. Elliott, A. Hahn, **M. Moe**)



- First time **enriched Ge** detector was used in experiment (ITEP-ErFI; **1987**)

## Main achievements in 1988-2001

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- $T_{1/2}(0\nu; {}^{76}\text{Ge}) > (1.6-1.9) \cdot 10^{25} \text{ y}$ ;  
(HM and IGEX; enriched HPGe detectors)
- $T_{1/2}(0\nu) > 10^{22}-10^{23} \text{ y}$  for  ${}^{136}\text{Xe}$ ,  ${}^{82}\text{Se}$ ,  
 ${}^{116}\text{Cd}$ ,  ${}^{100}\text{Mo}$
- $2\nu$ -decay was detected for many nuclei  
(TPC, ELEGANT-V, NEMO-2, HM, IGEX,  
Solotvino, Liq. Ar....) + transition to the  $0^+$   
excited states (Soudan, Modane, TUNL-  
ITEP)

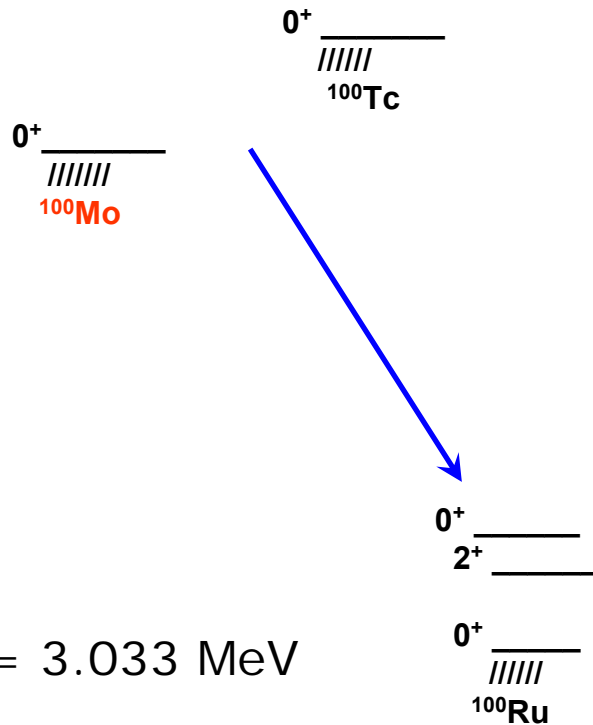
## II. PRESENT STATUS

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- 1. Introduction
- 2. Current experiments
  - **NEMO-3** and **CUORICINO**
  - “small-scale” experiments

# 1. Introduction

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There are 35 candidates for  $2\beta^-$ -decay

$$W \sim Q^5 (0\nu); W \sim Q^7 (0\nu\chi^0)$$

$$W \sim Q^{11} (2\nu)$$

# Candidates with $Q_{2\beta} > 2 \text{ MeV}$

Nuclei	$Q_{2\beta}$ , keV	Abundance, %
1. $^{48}\text{Ca}$	4272	0.187
2. $^{150}\text{Nd}$	3371.4	5.6
3. $^{96}\text{Zr}$	3350	2.8
4. $^{100}\text{Mo}$	3034.4	9.63
5. $^{82}\text{Se}$	2996	8.73
6. $^{116}\text{Cd}$	2805	7.49
7. $^{130}\text{Te}$	2527.5	<u>34.08</u>
8. $^{136}\text{Xe}$	2458.7	8.87
9. $^{124}\text{Sn}$	2287	5.79
10. $^{76}\text{Ge}$	2039.0	7.61
11. $^{110}\text{Pd}$	2000	11.72

Natural  $\gamma$ -rays background -  $E < 2.615 \text{ MeV}$ .  
So, there are **6 gold** and **5 silver** isotopes

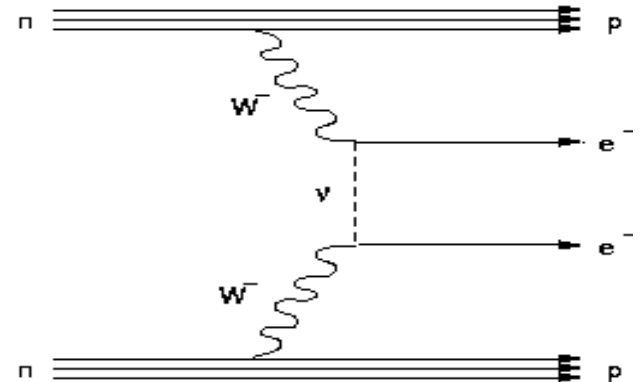
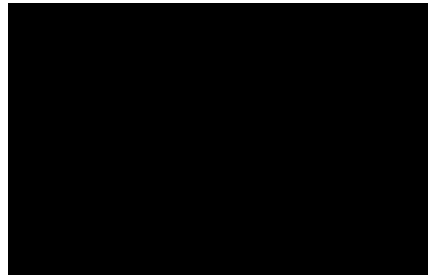
# NEUTRINOLESS DOUBLE BETA DECAY

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Experimental  
signature:

2 electrons

$$E_{\beta 1} + E_{\beta 2} = Q_{\beta\beta}$$



## Oscillation experiments $\Rightarrow$ Neutrino is massive!!!

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- However, the oscillatory experiments cannot solve the problem of the origin of neutrino mass (**Dirac or Majorana?**) and cannot provide information about the absolute value of mass (because the  $\Delta m^2$  is measured).
- This information can be obtained in  $2\beta$ -decay experiments.

$$\langle m_\nu \rangle = \left| \sum |U_{ej}|^2 e^{i\phi_j} m_j \right|$$

Thus searches for double beta decay are sensitive not only to masses but also to mixing elements and phases  $\phi_j$ .

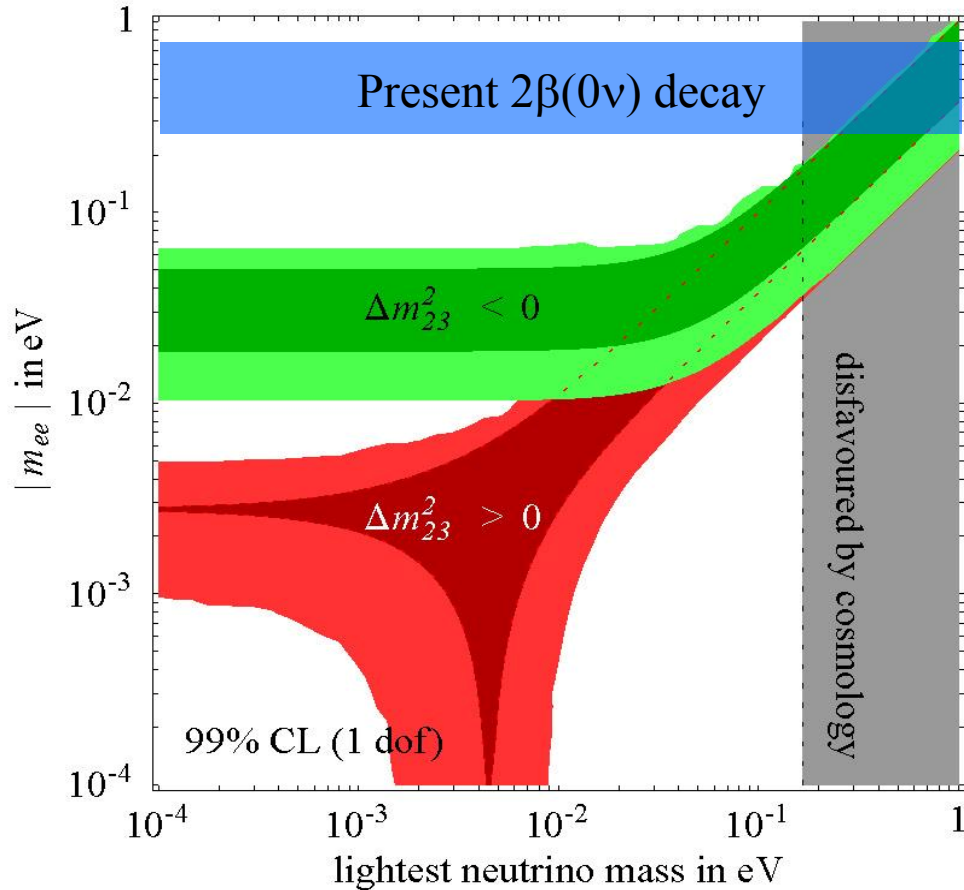


## What one can extract from $2\beta$ -decay experiments? $\Rightarrow$

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- Nature of neutrino mass (**Dirac or Majorana?**).
- Absolute mass scale (value or limit on  $m_1$ ).
- Type of hierarchy (normal, inverted, quasi-degenerated).
- **CP** violation in the lepton sector.

# DBD and neutrino mass hierarchy



**Degenerate:** can be tested

**Inverted:** can be tested by next generation of  $2\beta$  experiments.

**Normal:** inaccessible (new approach is needed)

$\beta$ :  $m_\nu < 2$  eV

$2\beta$ :  $\langle m_\nu \rangle < 0.75$  eV

**Cosmology :**  $\Sigma m_\nu < 0.2-0.6$  eV

( $\sim 0.04$  eV)

# Best present limits on $\langle m_\nu \rangle$

Nuclei	$T_{1/2}, y$	$\langle m_\nu \rangle, eV$ QRPA	$\langle m_\nu \rangle, eV$ [SM]	Experiment
$^{76}\text{Ge}$	$>1.9 \cdot 10^{25}$ $\approx 1.2 \cdot 10^{25} (?)$ $\approx 2.2 \cdot 10^{25} (?)$	$< 0.22-0.41$ $\approx 0.28-0.52 (?)$ $\approx 0.21-0.38 (?)$	$< 0.69$ $\approx 0.87 (?)$ $\approx 0.64 (?)$	HM Part of HM'04 Part of HM'06
	$>1.6 \cdot 10^{25}$	$<0.24-0.44$	$<0.75$	IGEX
$^{130}\text{Te}$	$>2.8 \cdot 10^{24}$	$< 0.35-0.59$	$< 0.77$	CUORICINO
$^{100}\text{Mo}$	$>1.1 \cdot 10^{24}$	$< 0.45-0.93$	-	NEMO
$^{136}\text{Xe}$	$>4.5 \cdot 10^{23}$	$< 1.41-2.67$	$< 2.2$	DAMA
$^{82}\text{Se}$	$>3.6 \cdot 10^{23}$	$< 0.89-1.61$	$< 2.3$	NEMO
$^{116}\text{Cd}$	$>1.7 \cdot 10^{23}$	$< 1.45-2.76$	$< 1.8$	SOLOTVINO

# A Recent Claim

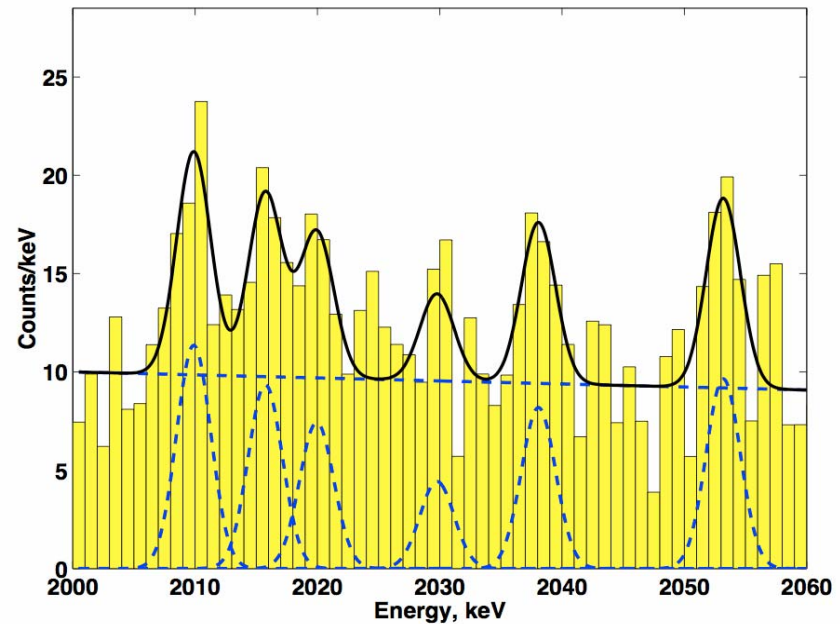
Klapdor-Kleingrothaus H V, Krivosheina I V, Dietz A and Chkvorets O, *Phys. Lett. B* **586** 198 (2004).

Used five  $^{76}\text{Ge}$  crystals, with a total of 10.96 kg of mass, and 71 kg-years of data

$$\tau_{1/2} = 1.2 \times 10^{25} \text{ y} \quad (4.2 \sigma)$$

$$0.24 < m_\nu < 0.58 \text{ eV} \quad (\pm 3 \text{ sigma})$$

(NME from Eur. Lett. 13(1990)31)



There are some problems with this result:

- 1) Only one measurement.
- 2) Only  $\sim 4\sigma$  level (independent analysis gives even  $\sim 2.7\sigma$ ).
- 3) In contradiction with HM'01 and IGEX.
- 4) Moscow part of Collaboration: **NO EVIDENCE.**
- 5)  $^{214}\text{Bi}$  peaks are overestimated.
- 6) "Total" and "analyzed" spectra are not the same.
- 7) Chkvorets'08 –  $1.3\sigma$

" $2\beta$  community": very conservative reaction

In any case new experiments are needed, which will confirm (or reject) this result

Mod.Phys.Lett. A21(2006)1547

Old data, new pulse shape anal.

$$\tau_{1/2} = 2.23^{+0.44}_{-0.31} \times 10^{25} \text{ y} \quad (6 \sigma)$$

$$m_\nu = 0.32 \pm 0.03 \text{ eV}$$

$$n = 11 \pm 1.8 \text{ events} \Rightarrow$$

where is a statistical error?!

non-correct peak position?!

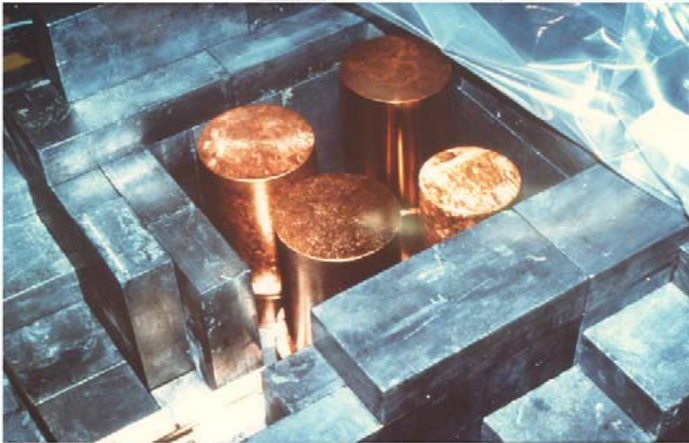
# Heidelberg-Moscow experiment

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**Gran Sasso**  
5 HPGe detectors  
(~ 11 kg of  $^{76}\text{Ge}$ )

**1990-2003**  
(full statistics:  
71.7 kg·y)



# Two neutrino double beta decay

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- Second order of weak interaction
- Direct measurement of NME values!  
⇒
  - The only possibility to check the quality of NME calculations!!!
  - $g_{pp}$  (QRPA parameter  $\Rightarrow$  NME( $0\nu$ !))  
⇓
- This is why it is very important to measure this type of decay for many nuclei, for different processes ( $2\beta^-$ ,  $2\beta^+$ ,  $K\beta^+$ ,  $2K$ , excited states) and with high accuracy.



# Two neutrino double beta decay

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- By present time  $2\beta(2\nu)$  decay was detected in **10** nuclei:  
 $^{48}\text{Ca}$ ,  $^{76}\text{Ge}$ ,  $^{82}\text{Se}$ ,  $^{96}\text{Zr}$ ,  $^{100}\text{Mo}$ ,  $^{116}\text{Cd}$ ,  $^{128}\text{Te}$ ,  $^{130}\text{Te}$ ,  $^{150}\text{Nd}$ ,  
 $^{238}\text{U}$

For  $^{100}\text{Mo}$  and  $^{150}\text{Nd}$   $2\beta(2\nu)$  transition to  $0^+$  **excited states** was detected too

**ECEC(2ν)** in  $^{130}\text{Ba}$  was detected in geochemical experiment

Main goal is: precise investigation of this decay

## Recommended values for half-lives:

- $^{48}\text{Ca}$  -  $(4.4^{+0.6}_{-0.5}) \cdot 10^{19}$  y
  - $^{76}\text{Ge}$  -  $(1.5 \pm 0.1) \cdot 10^{21}$  y
  - $^{82}\text{Se}$  -  $(0.92 \pm 0.07) \cdot 10^{20}$  y
  - $^{96}\text{Zr}$  -  $(2.3 \pm 0.2) \cdot 10^{19}$  y
  - $^{100}\text{Mo}$  -  $(7.1 \pm 0.4) \cdot 10^{18}$  y
  - $^{100}\text{Mo}$  -  $^{100}\text{Ru} (0^+_{1})$  -  
 $(5.9^{+0.8}_{-0.6}) \cdot 10^{20}$  y
  - $^{116}\text{Cd}$  -  $(2.8 \pm 0.2) \cdot 10^{19}$  y
  - $^{128}\text{Te}(\text{geo})$  -  $(1.9 \pm 0.4) \cdot 10^{24}$  y
  - $^{130}\text{Te}$  -  $(6.8^{+1.2}_{-1.1}) \cdot 10^{20}$  y
  - $^{150}\text{Nd}$  -  $(8.2 \pm 0.9) \cdot 10^{18}$  y
  - $^{150}\text{Nd}$  -  $^{150}\text{Sm} (0^+_{1})$  -  
 $(1.33^{+0.45}_{-0.26}) \cdot 10^{20}$  y
  - $^{238}\text{U}(\text{rad})$  -  $(2.0 \pm 0.6) \cdot 10^{21}$  y
- ECEC(2v):
- $^{130}\text{Ba}(\text{geo})$  -  $(2.2 \pm 0.5) \cdot 10^{21}$  y



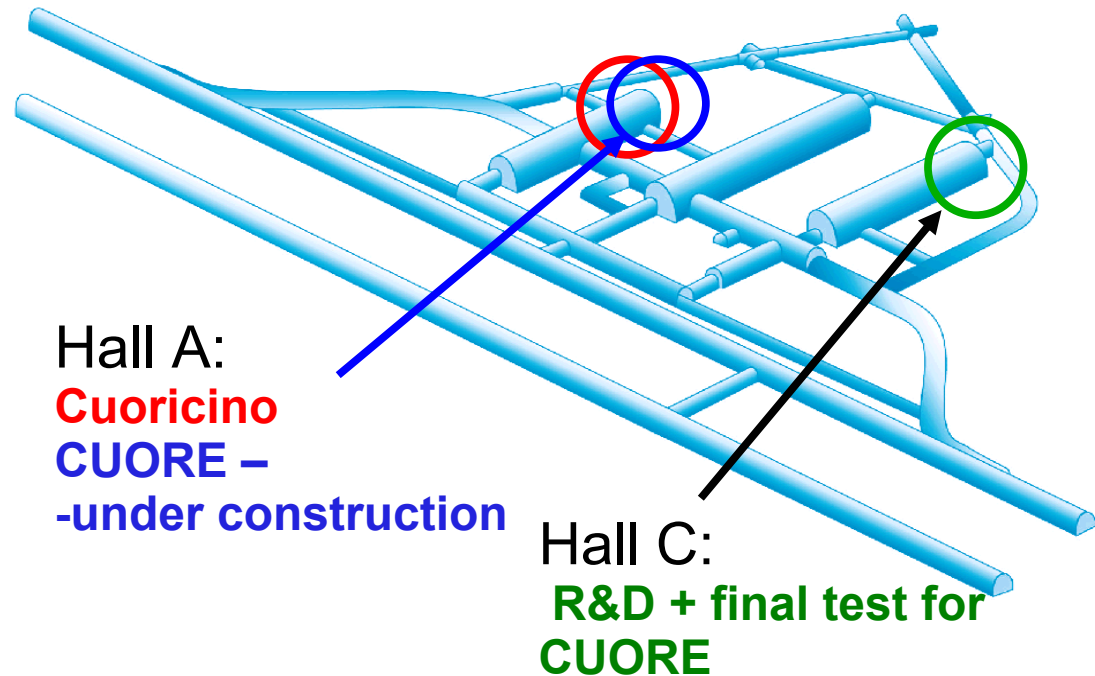
## 2. CURRENT EXPERIMENTS

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- **NEMO-3** and **CUORICINO**
- Others (TGV, Baksan, DAMA, COBRA, ITEP-TPC, TUNL-ITEP, excited states,...)

# CUORICINO

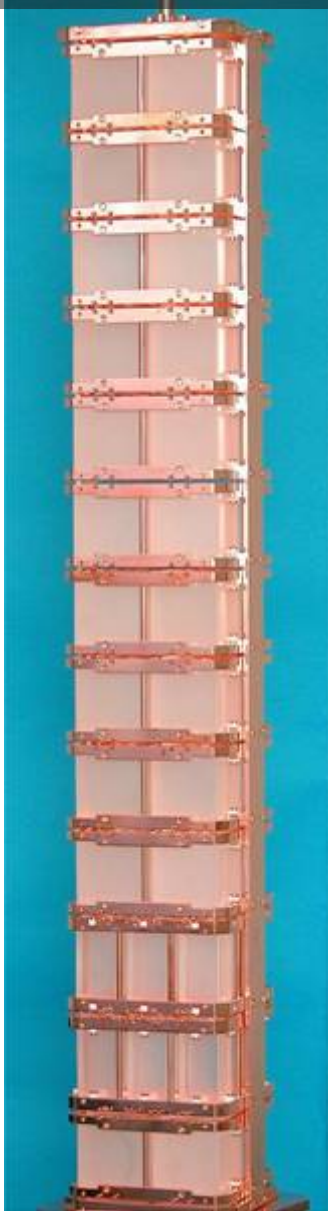
INFN - Laboratori Nazionali del Gran Sasso - L'Aquila – Italy



3200 m.w.e overburden - cosmic rays are no more a bkg problem

- ★ n flux is reduced to  $\sim 10^{-6}$  n/cm<sup>2</sup>/s
- ★  $\mu$  flux is  $\sim 2$ /m<sup>2</sup>/h

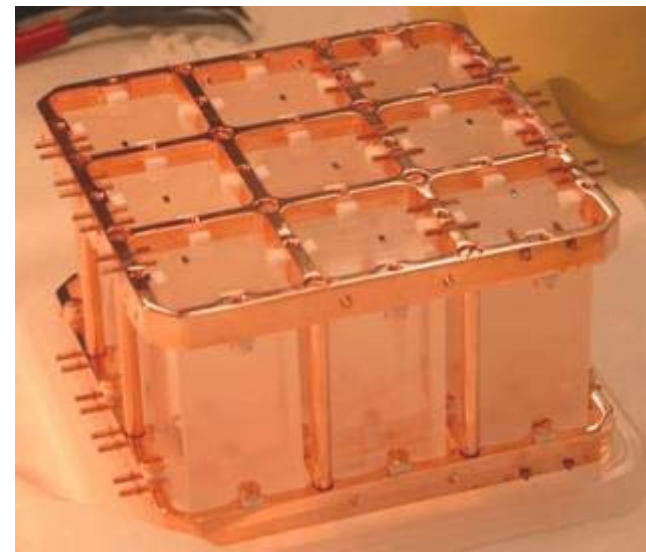
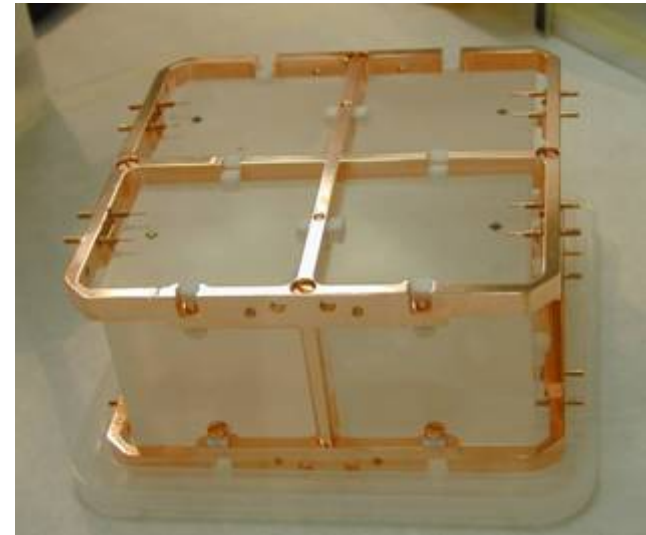
# Cuoricino



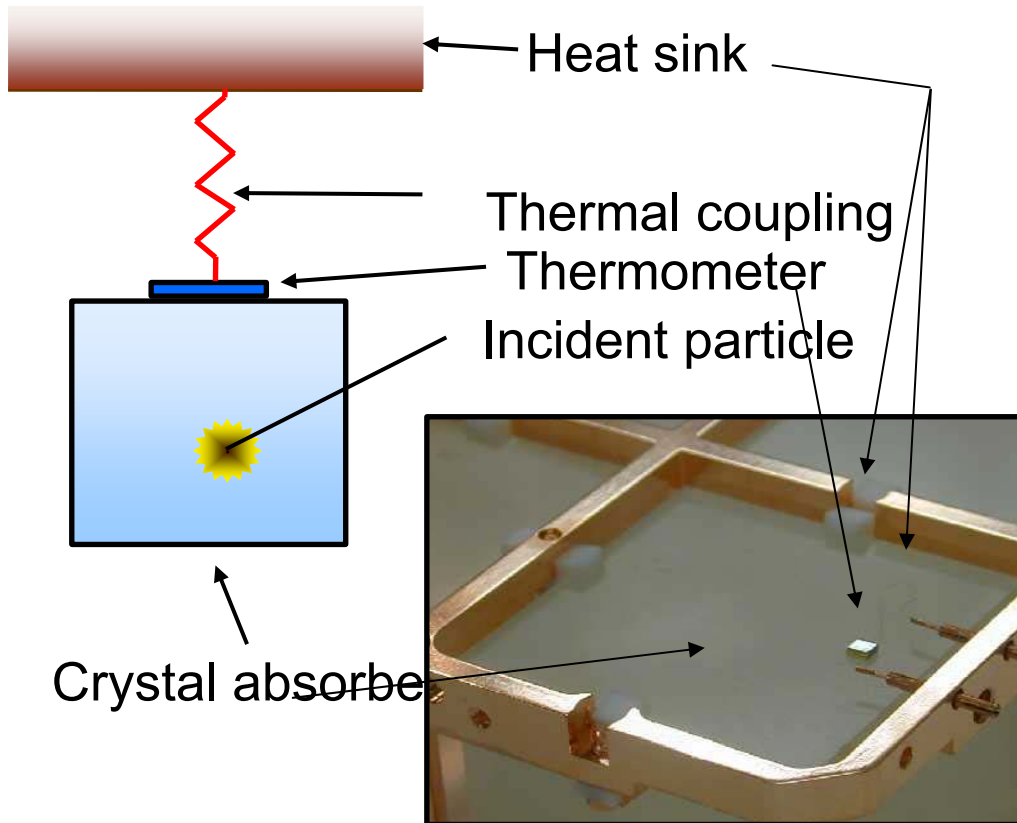
11 modules  
4 detectors each  
Dimension: 5x5x5 cm<sup>3</sup>  
Mass: 790 g

Total mass  
40.7 kg  
(~11 kg of <sup>130</sup>Te)

2 modules  
9 detectors each,  
Dimension: 3x3x6 cm<sup>3</sup>  
Mass: 330 g



# Low Temperature Detectors (LTD)



## Detection Principle

$$\Delta T = E/C$$

$C$ : thermal capacity

low  $C$

low  $T$  (i.e.  $T \ll 1\text{K}$ )

dielectrics, superconductors

ultimate limit to  $E$  resolution:  
statistical fluctuation of internal

energy  $U$

$$\langle \Delta U^2 \rangle = k_B T^2 C$$

## Thermal Detectors Properties

good energy resolution

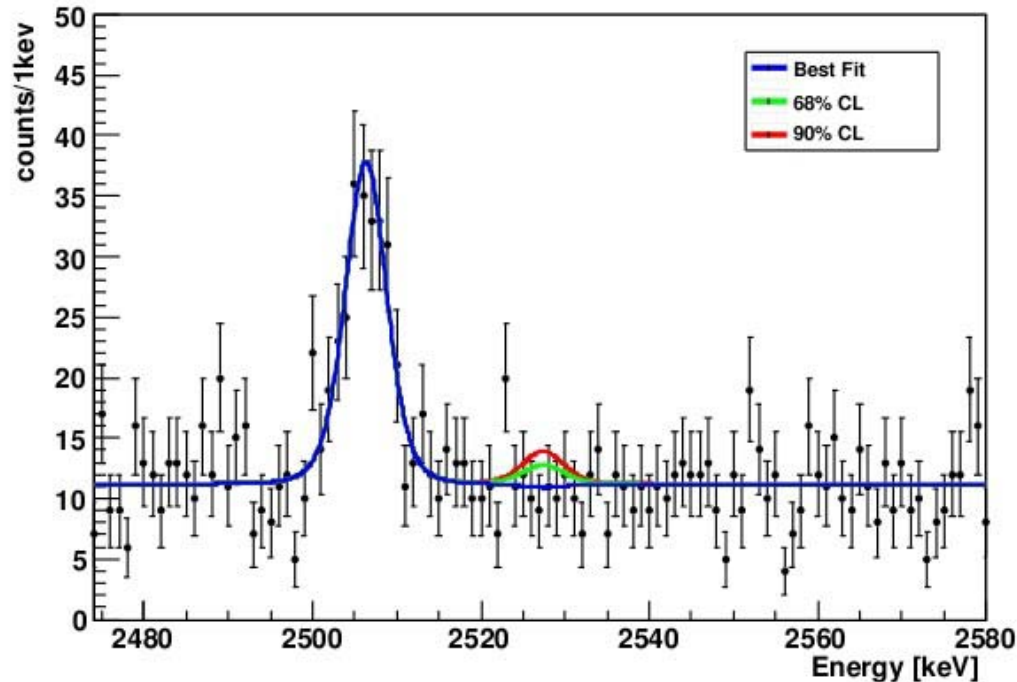
wide choice of absorber materials

true calorimeters

slow  $\tau = C/G \sim 1 \div 10^3$  ms

**$T = 8$  mK**

# Cuoricino result on $^{130}\text{Te}$ $\beta\beta_{0\nu}$ decay



Total statistic  
 $\sim 19.75 \text{ kg } (^{130}\text{Te}) \times y$

$$b = 0.18 \pm 0.01 \text{ c/keV/kg/y}$$

Maximum Likelihood  
 flat background + fit of 2505 peak

*Anticoincidence background spectrum the bb-0n region*

$$\tau_{1/2}^{0\nu} \geq 2.8 \cdot 10^{24} \text{ y (90\% CL)}$$



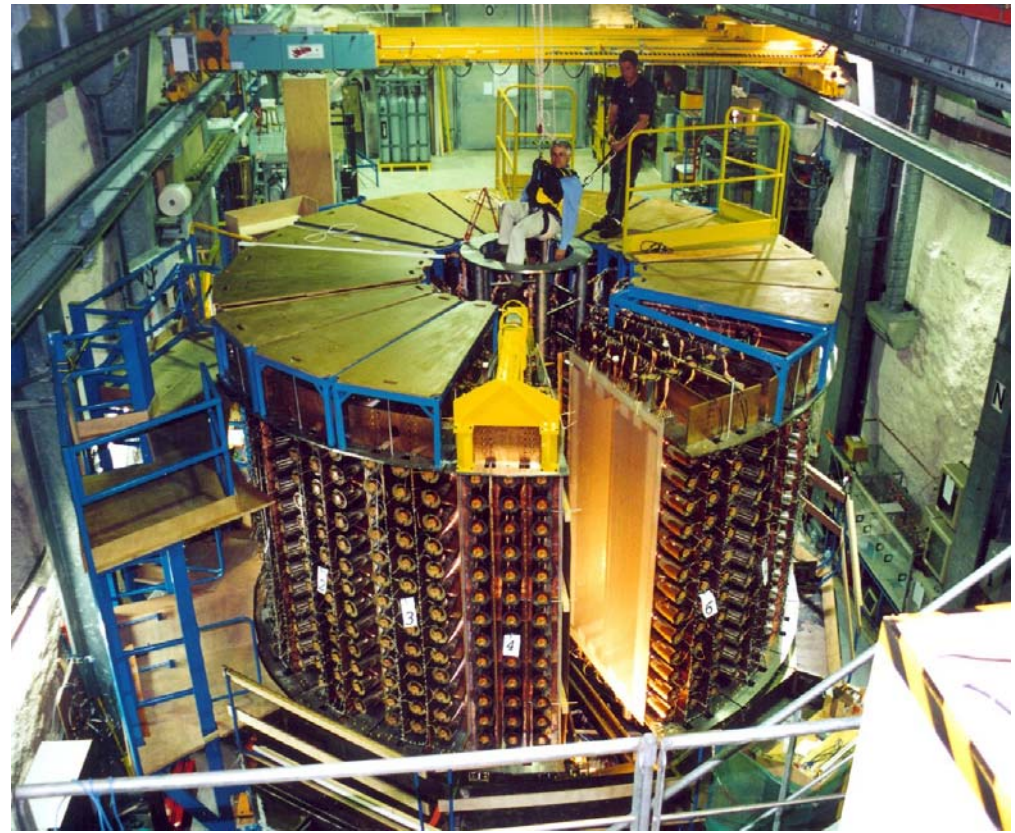
$$\langle m_{\nu} \rangle \leq 0.3 - 0.7 \text{ eV (90\% CL)}$$

[Experiment is stopped in July 2008]

# NEMO-3 Collaboration

(Neutrino Ettore Majorana Observatory)

60 physicists, 17 labs





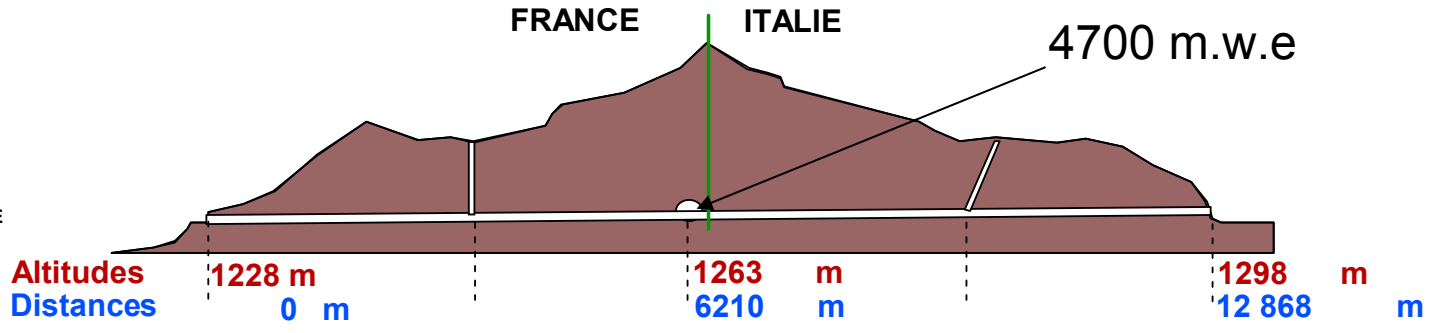
# Laboratoire Souterrain de Modane



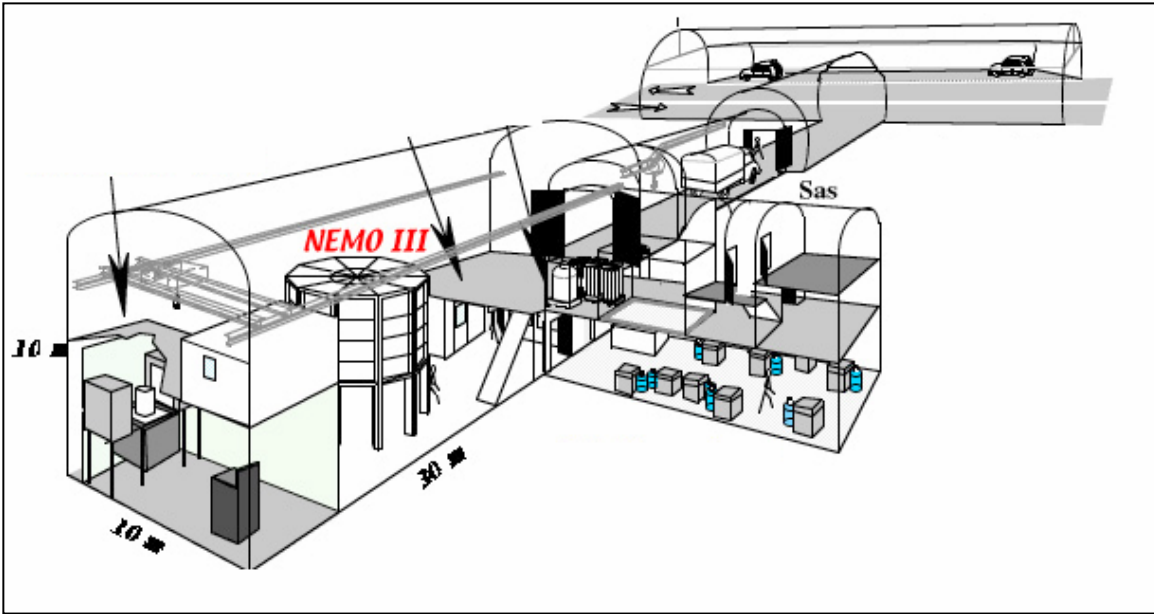
COMMISSARIAT À L'ÉNERGIE ATOMIQUE



DIRECTION DES SCIENCES DE LA MATIÈRE

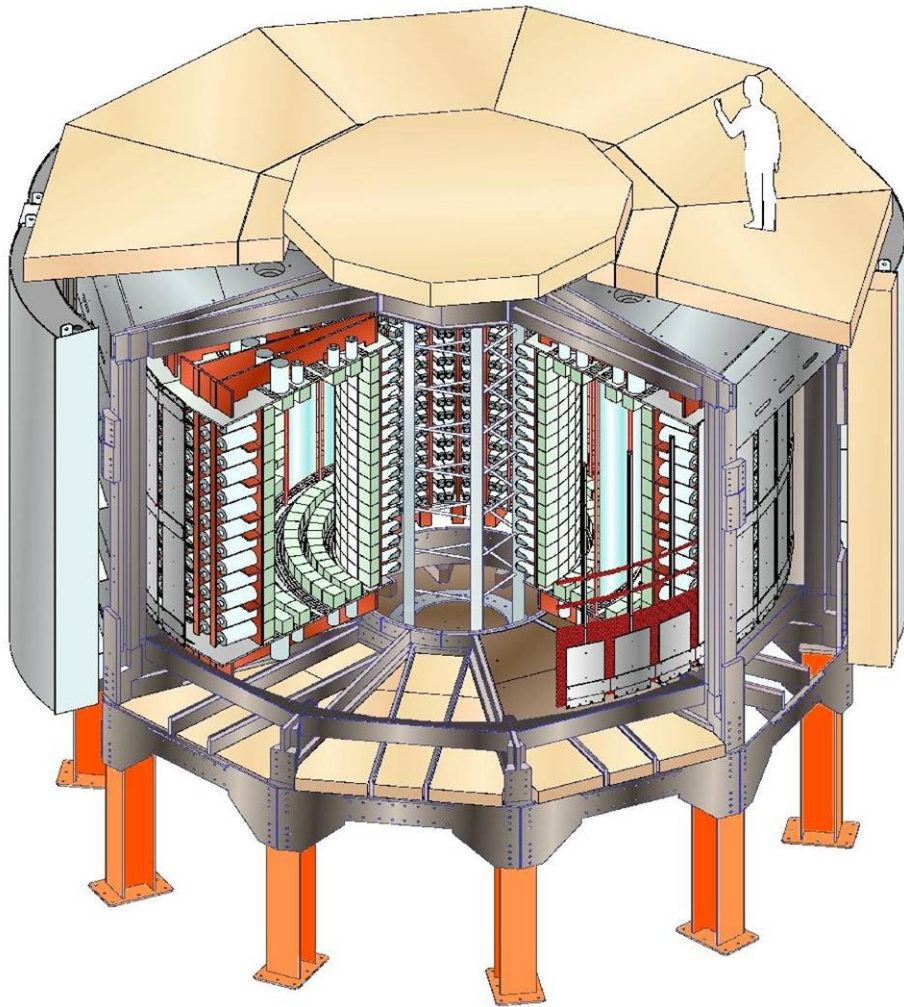


Built for  $\tau$  experiment (proton decay) in 1981-1982



# The NEMO3 detector

Fréjus Underground Laboratory : 4800 m.w.e.



Source: 10 kg of  $\beta\beta$  isotopes  
cylindrical,  $S = 20 \text{ m}^2$ ,  $60 \text{ mg/cm}^2$

Tracking detector:

drift wire chamber operating  
in Geiger mode (6180 cells)

Gas: He + 4% ethyl alcohol + 1% Ar + 0.1% H<sub>2</sub>O

Calorimeter:

1940 plastic scintillators  
coupled to low radioactivity PMTs

**Magnetic field: 25 Gauss**

**Gamma shield: Pure Iron (18 cm)**

**Neutron shield: borated water (~30 cm) + Wood (Top/Bottom/Gapes between water tanks)**



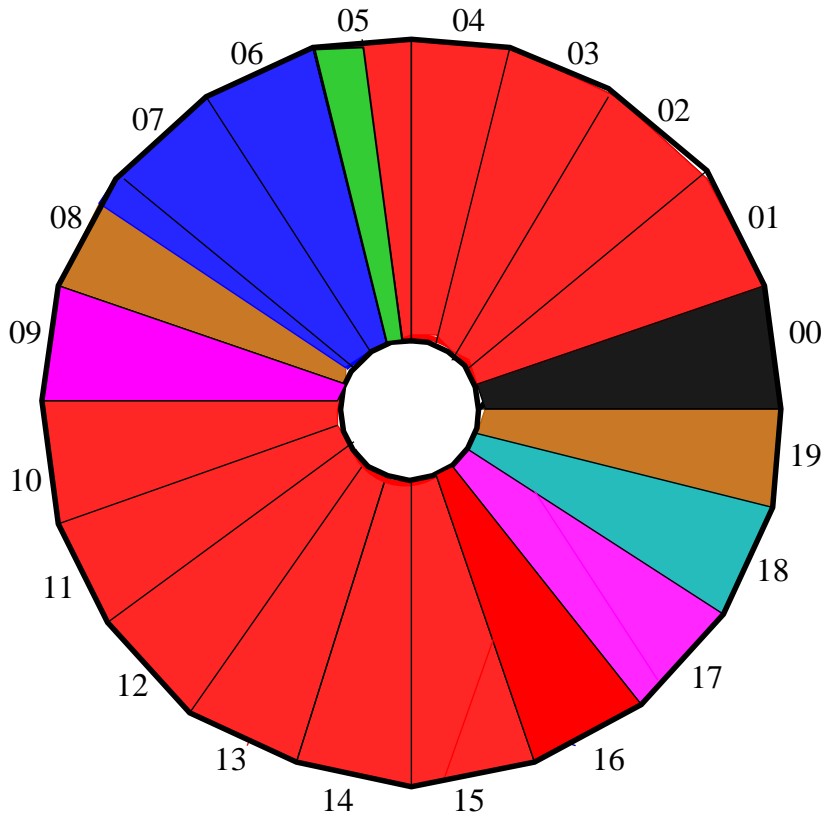
**Able to identify  $e^-$ ,  $e^+$ ,  $\gamma$  and  $\alpha$**



# Finished detector



# $\beta\beta$ decay isotopes in NEMO-3 detector



**$^{100}\text{Mo}$  6.914 kg**  $Q_{\beta\beta} = 3034 \text{ keV}$   
 **$^{82}\text{Se}$  0.932 kg**  $Q_{\beta\beta} = 2995 \text{ keV}$

$\beta\beta 0\nu$  search

$\beta\beta 2\nu$  measurement

**$^{116}\text{Cd}$  405 g**  
 $Q_{\beta\beta} = 2805 \text{ keV}$

**$^{96}\text{Zr}$  9.4 g**  
 $Q_{\beta\beta} = 3350 \text{ keV}$

**$^{150}\text{Nd}$  37.0 g**  
 $Q_{\beta\beta} = 3367 \text{ keV}$

**$^{48}\text{Ca}$  7.0 g**  
 $Q_{\beta\beta} = 4272 \text{ keV}$

**$^{130}\text{Te}$  454 g**  
 $Q_{\beta\beta} = 2529 \text{ keV}$

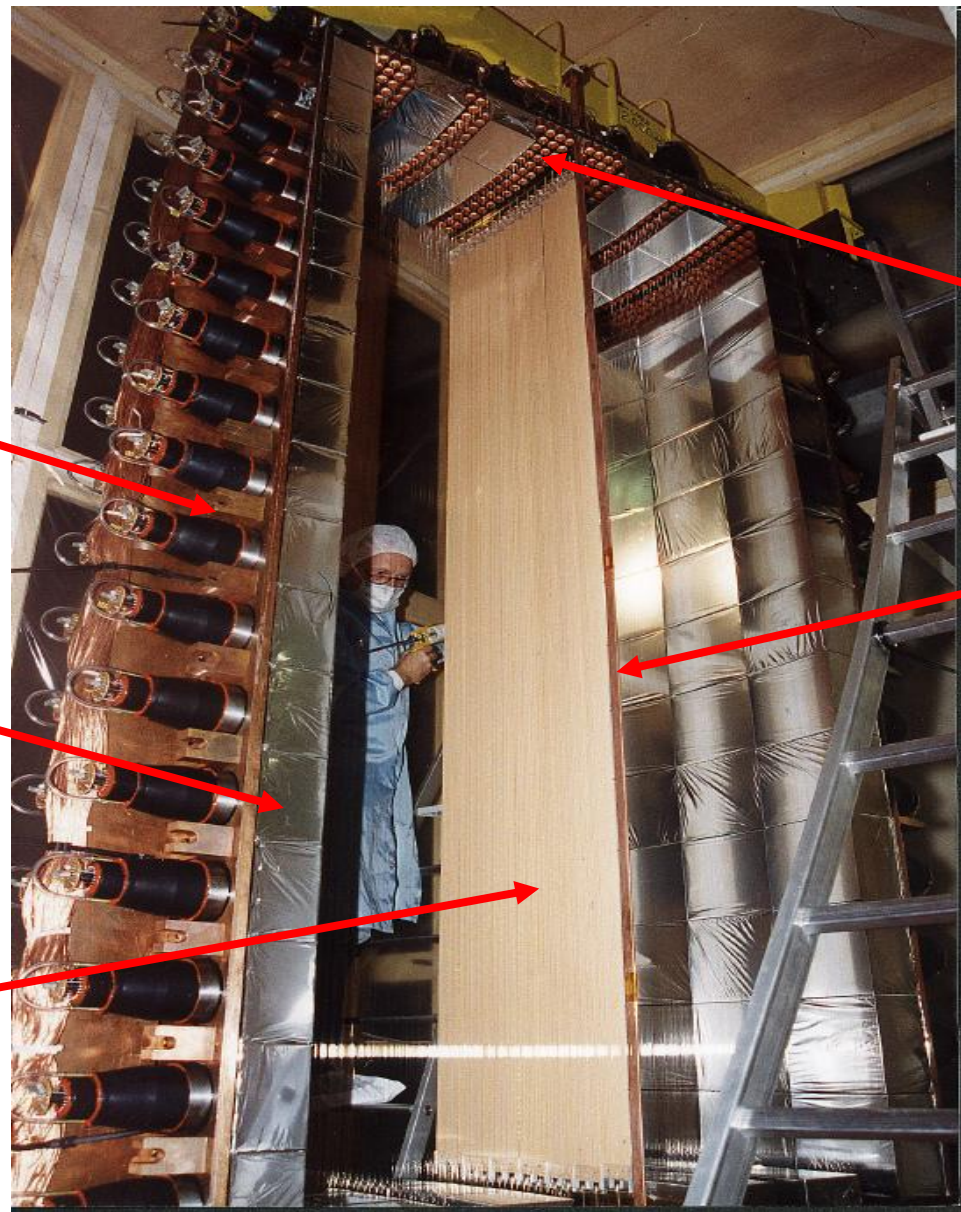
**$^{\text{nat}}\text{Te}$  491 g**

**Cu 621 g**

External bkg measurement

(All enriched isotopes produced in Russia)

# Sector interior view



PMT

scintillators

$\beta\beta$  isotope foils

cathode rings  
wire chamber

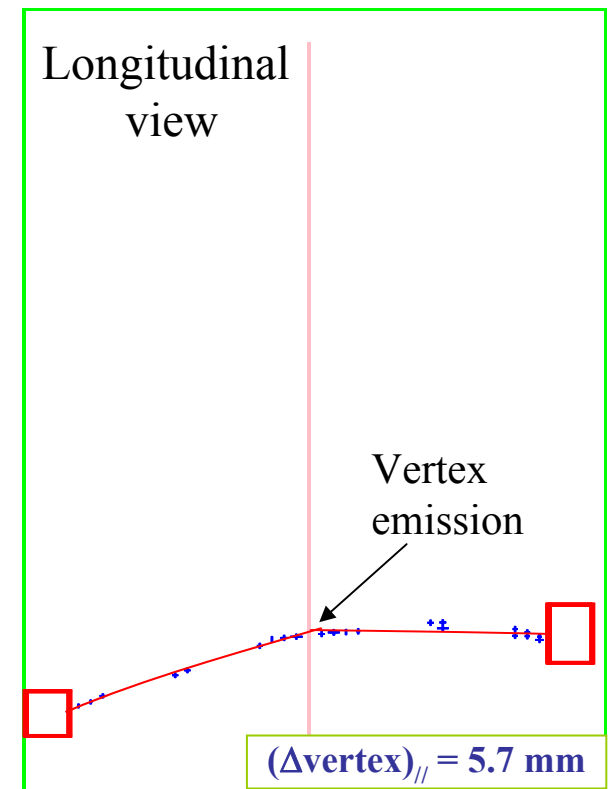
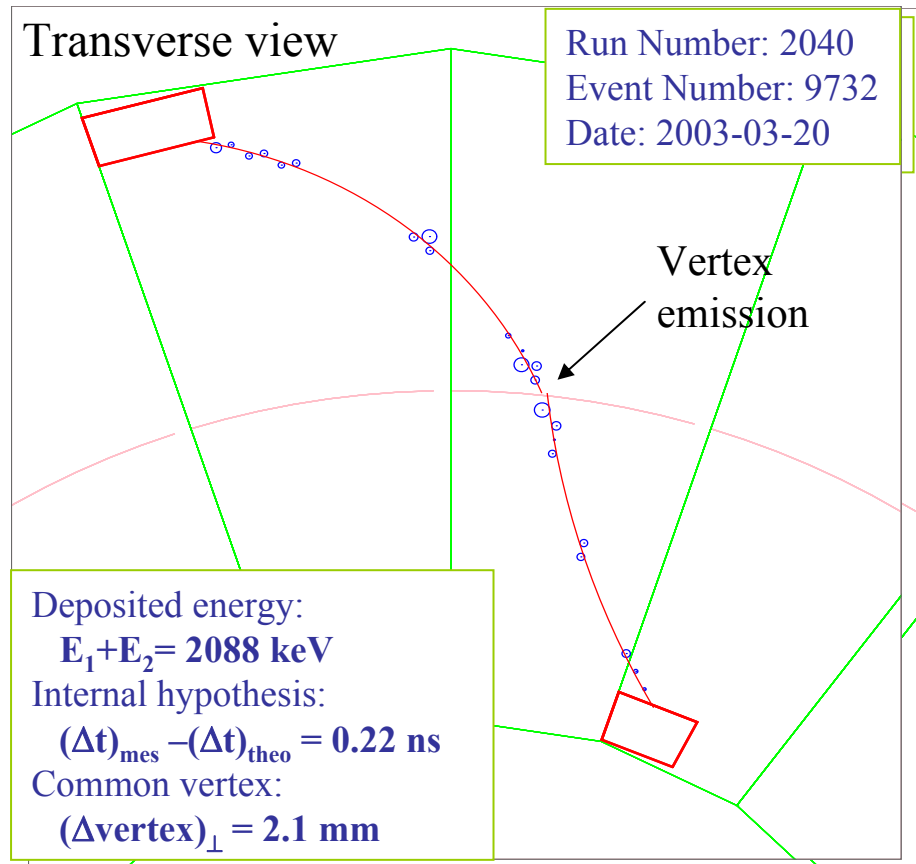
calibration tube

Calibration source

- $^{207}\text{Bi}$
- $2e^-$  (IC) lines  
 $\sim 0.5$  ,  $\sim 1$  MeV
- $^{90}\text{Sr}$
- $^{60}\text{Co}$

# $\beta\beta$ events selection in NEMO-3

Typical  $\beta\beta 2\nu$  event observed from  $^{100}\text{Mo}$



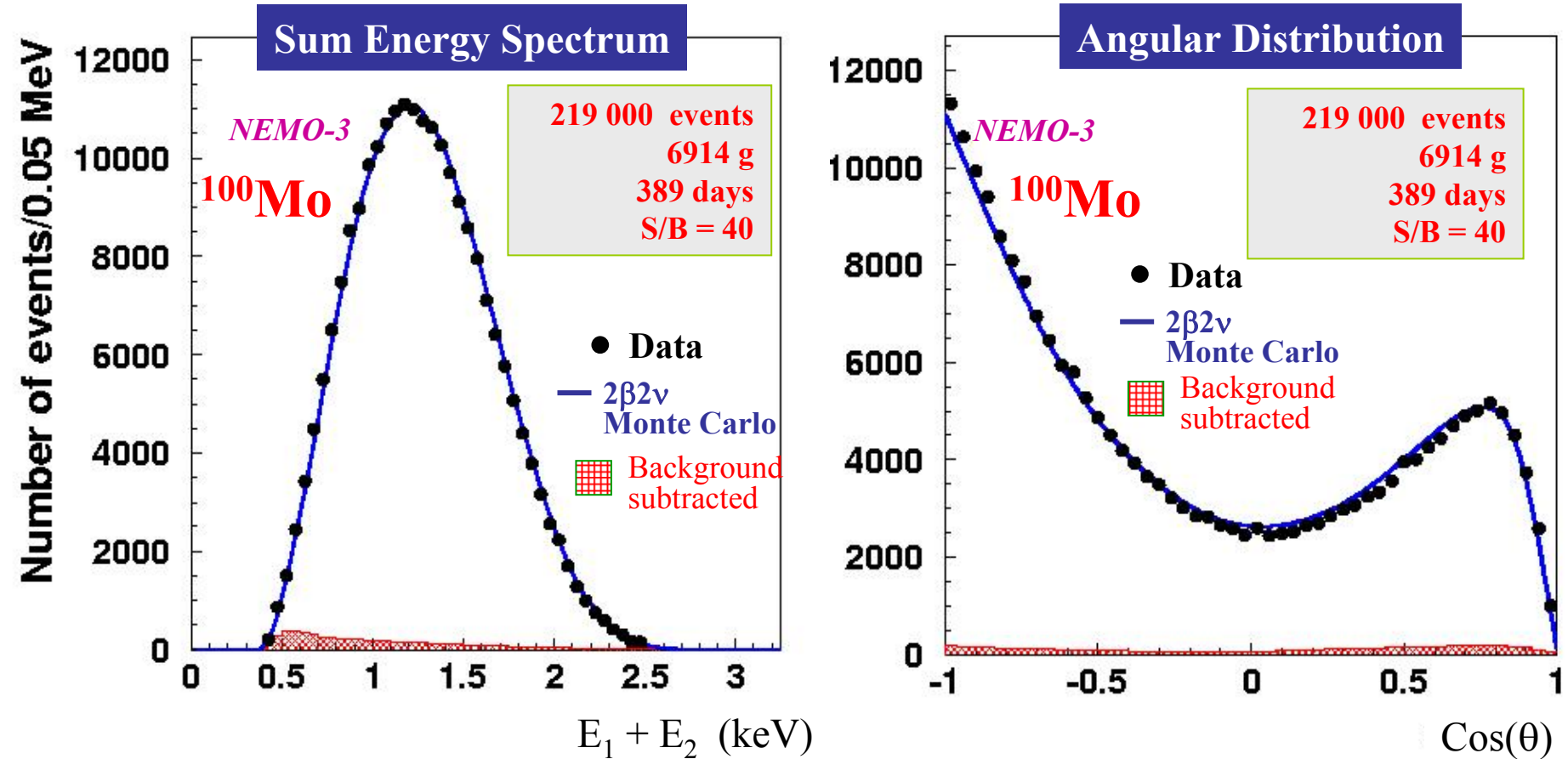
**Trigger:** at least 1 PMT > 150 keV  
 $\geq 3$  Geiger hits (2 neighbour layers + 1)

Trigger rate = 7 Hz

$\beta\beta$  events: 1 event every 2.5 minutes

# $^{100}\text{Mo}$ $2\beta 2\nu$ preliminary results

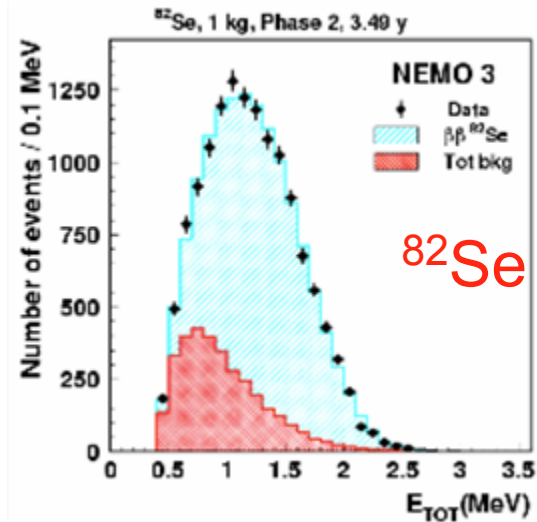
(Data Feb. 2003 – Dec. 2004)



7.37 kg.y

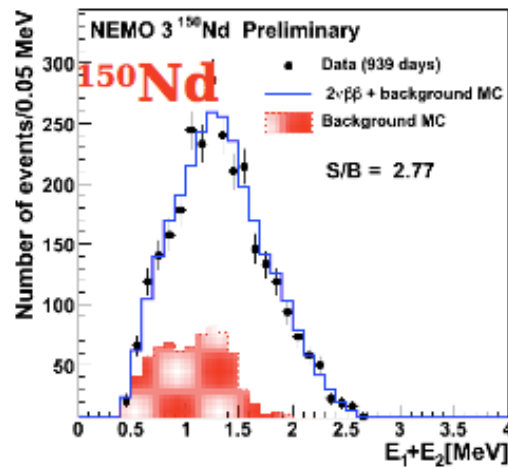
$T_{1/2} = 7.11 \pm 0.02$  (stat)  $\pm 0.54$  (syst)  $\times 10^{18}$  y

# 2νββ results for other isotopes



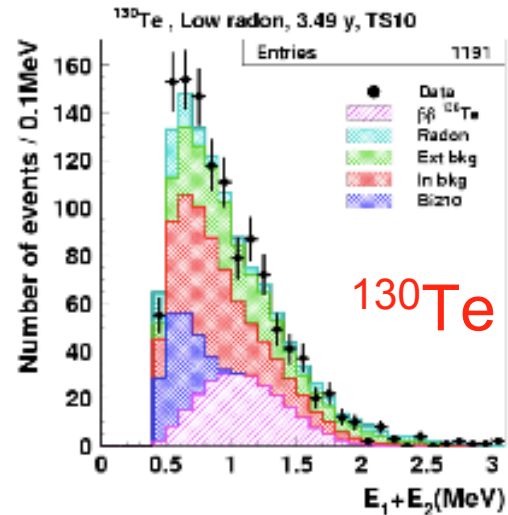
$$[9.6 \pm 0.1(\text{stat}) \pm 1.0(\text{sys})] \times 10^{19} \text{ yr}$$

$$M^{2\nu} = 0.049 \pm 0.004$$



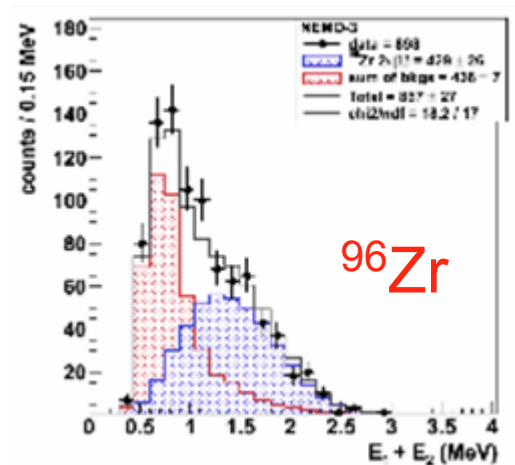
$$[9.11 \pm 0.25(\text{stat}) \pm 0.63(\text{sys})] \times 10^{18} \text{ yr}$$

$$M^{2\nu} = 0.030 \pm 0.002$$



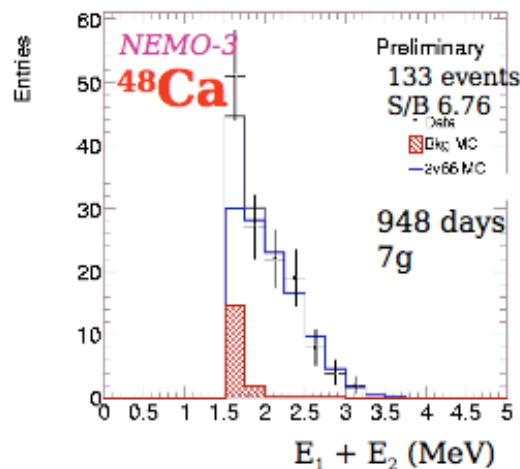
$$[7.0^{+1.0}_{-0.8}(\text{stat})^{+1.1}_{-0.9}(\text{sys})] \times 10^{20} \text{ yr}$$

$$M^{2\nu} = 0.0173 \pm 0.0025$$



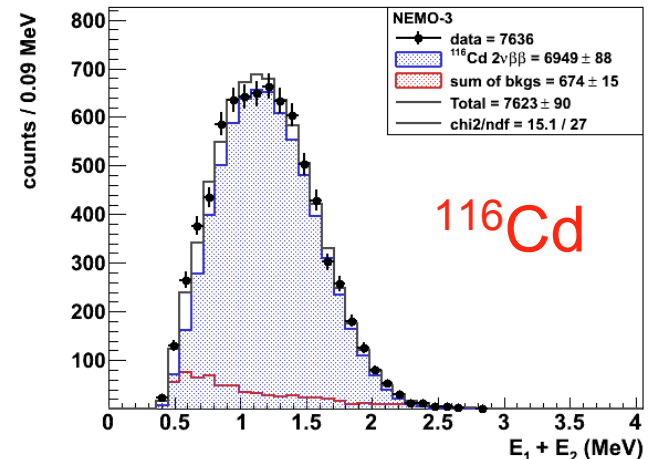
$$[2.35 \pm 0.14(\text{stat}) \pm 0.16(\text{sys})] \times 10^{19} \text{ yr}$$

$$M^{2\nu} = 0.049 \pm 0.002$$



$$[4.4^{+0.5}_{-0.4}(\text{stat}) \pm 0.4(\text{sys})] \times 10^{19} \text{ yr}$$

$$M^{2\nu} = 0.0238 \pm 0.0015$$



$$[2.88 \pm 0.04(\text{stat}) \pm 0.16(\text{sys})] \times 10^{19} \text{ yr}$$

$$M^{2\nu} = 0.0685 \pm 0.0025$$

# Summary of $2\nu\beta\beta$ results

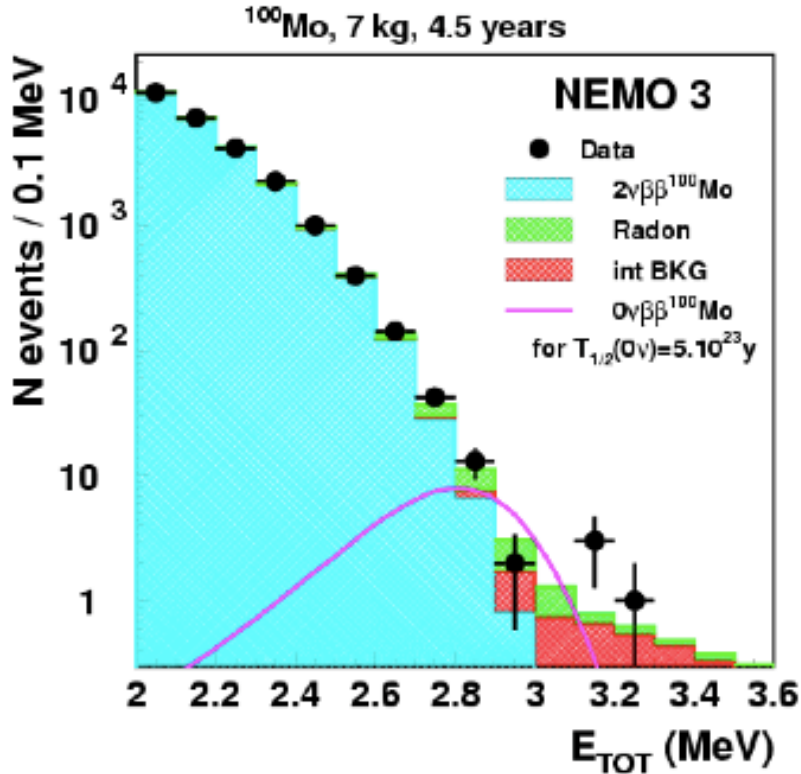
Isotope	S/B	$(2\nu\beta\beta), \gamma$
$^{100}\text{Mo}$	<b>40</b>	$(7.11 \pm 0.02(\text{stat}) \pm 0.54(\text{syst})) \cdot 10^{18}$ (SSD favoured) *
$^{100}\text{Mo}(0^+_{1})$	<b>3</b>	$(5.7^{+1.3}_{-0.9}(\text{stat}) \pm 0.8(\text{syst})) \cdot 10^{20}$ ** [NPA 781 (2006) 209]
$^{82}\text{Se}$	<b>4</b>	$(9.6 \pm 0.3(\text{stat}) \pm 1.0(\text{syst})) \cdot 10^{19}$ *
$^{116}\text{Cd}$	<b>7.5</b>	$(2.88 \pm 0.04(\text{stat}) \pm 0.16(\text{syst})) \cdot 10^{19}$ ***
$^{130}\text{Te}$	<b>0.35</b>	$(7.0^{+1.0}_{-0.8}(\text{stat}) +1.1_{-0.9}(\text{syst})) \cdot 10^{20}$ ***
$^{150}\text{Nd}$	<b>2.8</b>	$(9.11^{+0.25}_{-0.22}(\text{stat}) \pm 0.63(\text{syst})) \cdot 10^{18}$ *** [PRC 80 (2009) 032501R]
$^{96}\text{Zr}$	<b>1.0</b>	$(2.35 \pm 0.14(\text{stat}) \pm 0.16(\text{syst})) \cdot 10^{19}$ *** [NPA 847 (2010) 168]
$^{48}\text{Ca}$	<b>6.8</b>	$(4.4^{+0.5}_{-0.4}(\text{stat}) \pm 0.4(\text{syst})) \cdot 10^{19}$ ***

\* Phase 1 data, Phys. Rev. Lett. 95 (2005) 182302. Additional statistics are being analysed, to be published soon.

\*\* Phase 1 data.

\*\*\* Phases 1 and 2, preliminary.

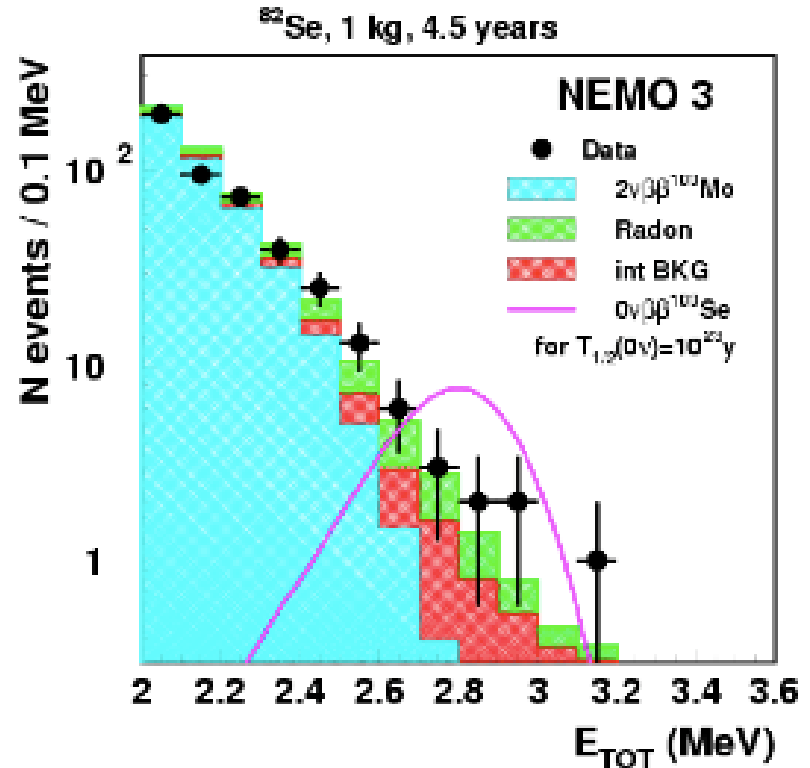
# $0\nu\beta\beta$ for $^{100}\text{Mo}$ (~7kg) and $^{82}\text{Se}$ (~1kg)



[2.8-3.2] MeV: DATA = 18; MC =  $16.4 \pm 1.4$

$T_{1/2}(0\nu) > 1.0 \times 10^{24}$  yr at 90%CL

$\langle m_\nu \rangle < (0.47 - 0.96)$  eV



[2.6-3.2] MeV: DATA = 14; MC =  $10.9 \pm 1.3$

$T_{1/2}(0\nu) > 3.2 \times 10^{23}$  yr at 90%CL

$\langle m_\nu \rangle < (0.94 - 2.5)$  eV



# Summary of $0\nu\beta\beta$ results

- No evidence for non conservation of the leptonic number
- Current limits on  $0\nu\beta\beta$  (at 90% C.L.):

Isotope	Exposure (kg·y)	$T_{1/2}(0\nu\beta\beta)$ , y	$\langle m_\nu \rangle$ , eV [NME ref.]
$^{100}\text{Mo}$	31	$> 1 \cdot 10^{24}$	$< 0.47 - 0.96$ [1-3]
$^{82}\text{Se}$	4.2	$> 3.2 \cdot 10^{23}$	$< 0.9 - 1.6$ [1-3]; $< 2.5$ [7]
$^{150}\text{Nd}$	0.095	$> 1.8 \cdot 10^{22}$	$< 1.7 - 2.4$ [4,5] ; $< 4.8 - 7.6$ [6]
$^{130}\text{Te}$	1.4	$> 9.8 \cdot 10^{22}$	$< 1.6 - 3.1$ [2,3]
$^{96}\text{Zr}$	0.031	$> 9.2 \cdot 10^{21}$	$< 7.2 - 19.5$ [2,3]
$^{48}\text{Ca}$	0.017	$> 1.3 \cdot 10^{22}$	$< 29.6$ [7]

- NME references:

- [1] M.Kortelainen and J.Suhonen, Phys.Rev. C 75 (2007) 051303(R)
- [2] M.Kortelainen and J.Suhonen, Phys.Rev. C 76 (2007) 024315
- [3] F.Simkovic, et al. Phys.Rev. C 77 (2008) 045503
- [4] V.A. Rodin et al. Nucl.Phys. A 793 (2007) 213
- [5] V.A. Rodin et al. Nucl.Phys. A 766(2006) 107
- [6] J.H.Hirsh et al. Nucl.Phys. A 582(1995) 124
- [7] E.Caurrier et al. Phys.Rev.Lett 100 (2008) 052503

## Most active “small” experiments

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- **TGV-II** (multi HPGe;  $^{106}\text{Cd}$ ; Modan)
- **TUNL-ITEP** (2xHPGe; excited states in  $^{100}\text{Mo}$ ,  $^{150}\text{Nd}$ ; USA)
- **Baksan** (proportional counter;  $^{136}\text{Xe}$ ,  $^{78}\text{Kr}$ )
- **DAMA-KIEV** (scintillators;  $^{136}\text{Ce}$ ,  $^{64}\text{Zn}$ ,  $^{180}\text{W}$ ...; Gran Sasso)
- **ITEP-Bordeaux** (HPGe; excited states:  $^{100}\text{Mo}$ ,  $^{82}\text{Se}$ ,  $^{150}\text{Nd}$ ,  $^{74}\text{Se}$ ,  $^{112}\text{Cd}$ ,...; Modan)
- **COBRA** (CdZnTe semiconductor; Gran Sasso)

# III. FUTURE EXPERIMENTS

---

- **Main goal is:**  
To reach a sensitivity  $\sim$  **0.01-0.1 eV** to  $\langle m_\nu \rangle$   
(inverted hierarchy region)
- **Strategy is:**
  - to investigate different isotopes (**>2-3**);
  - to use **different** experimental technique

**Here I have selected a few propositions which I believe will be realized in the nearest future (~3-10 years)**

---

- **CUORE** ( $^{130}\text{Te}$ , cryogenic thermal detector)
- **GERDA** ( $^{76}\text{Ge}$ , HPGe detector)
- **MAJORANA** ( $^{76}\text{Ge}$ , HPGe detector)
- **EXO** ( $^{136}\text{Xe}$ , TPC +  $\text{Ba}^+$ )
- **SuperNEMO** ( $^{82}\text{Se}$  or  $^{150}\text{Nd}$ , tracking detector)
- **KamLAND-Xe** ( $^{136}\text{Xe}$ , liquid scintillator)
- **SNO+** ( $^{150}\text{Nd}$ , liquid scintillator)

Other proposals: CANDLES, COBRA, XMASS, MOON, DCBA, NEXT, LUCIFER, ...

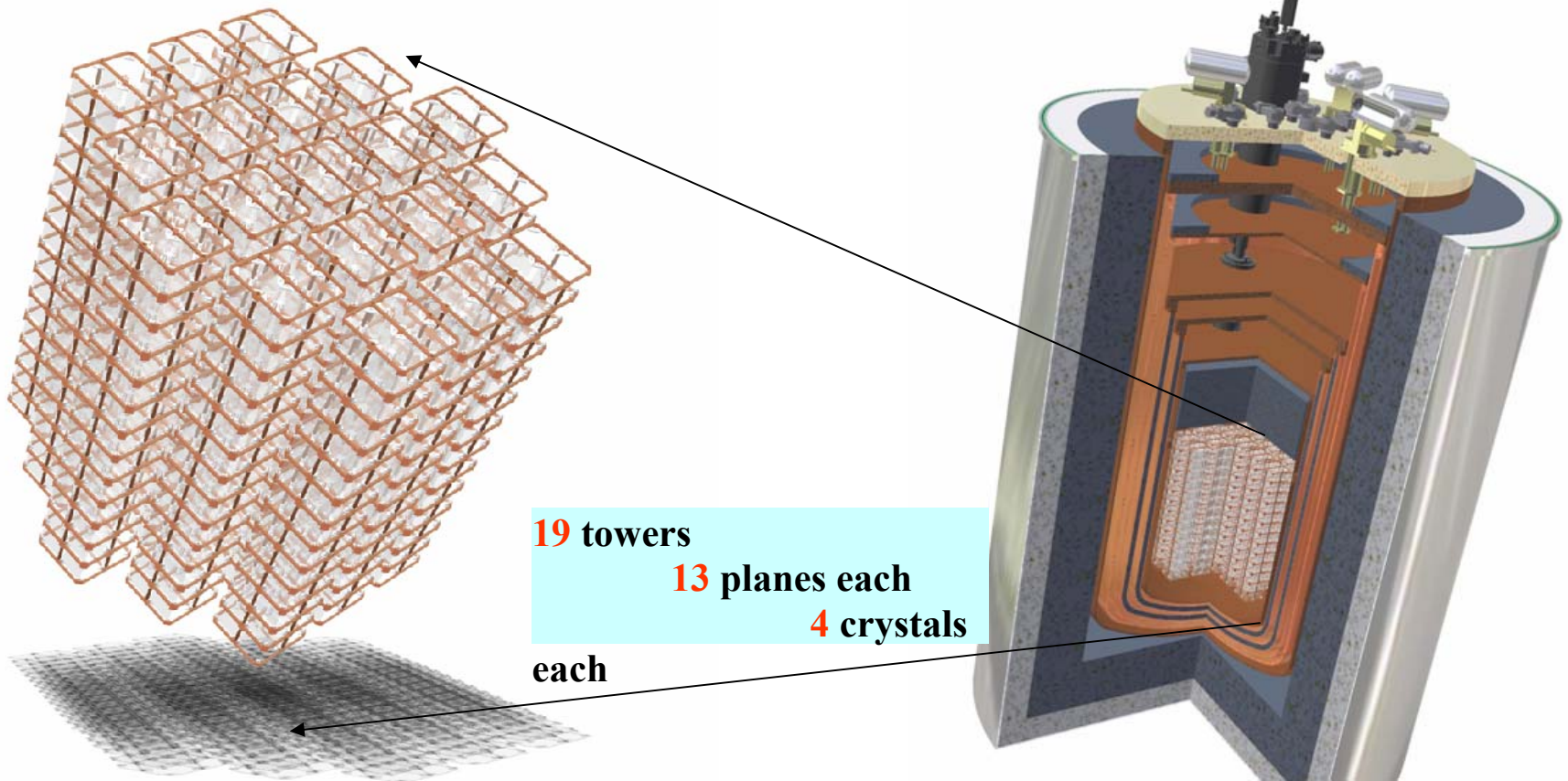
# SUMMARY TABLE

Experiment	Isotope	Mass, kg	$T_{1/2}$ , y	$\langle m_\nu \rangle$ , meV	Status
<b>CUORE</b>	$^{130}\text{Te}$	200	$2.1 \cdot 10^{26}$	40-90	Funded
<b>GERDA</b>	$^{76}\text{Ge}$	I. 17 II. 40 III. 1000	$3 \cdot 10^{25}$ $2 \cdot 10^{26}$ $6 \cdot 10^{27}$	70-200 10-40	Funded Funded R&D
<b>MAJORANA</b>	$^{76}\text{Ge}$	I. 30-60 II. 1000	$2 \cdot 10^{26}$ $6 \cdot 10^{27}$	70-200 10-40	Funded R&D
<b>EXO</b>	$^{136}\text{Xe}$	200 1000	$6.4 \cdot 10^{25}$ $8 \cdot 10^{26}$	100-200 30-60	Funded R&D
<b>SuperNEMO</b>	$^{82}\text{Se}$	100-200	$(1-2) \cdot 10^{26}$	40-100	R&D
<b>KamLAND-Xe</b>	$^{136}\text{Xe}$	400 1000	$\sim 4 \cdot 10^{26}$ $\sim 10^{27}$	40-80 25-50	Funded R&D
<b>SNO+</b>	$^{150}\text{Nd}$	56 500	$\sim 4.5 \cdot 10^{24}$ $\sim 3 \cdot 10^{25}$	100-300 40-120	Funded R&D

# CUORE

## Cryogenic Underground Observatory for Rare Events

- Closely packed array of 988 TeO<sub>2</sub> crystals 5×5×5 cm<sup>3</sup> (750 g)
  - 741 kg TeO<sub>2</sub> granular calorimeter
  - 600 kg Te = 203 kg <sup>130</sup>Te
- Single high granularity detector



# CUORE schedule



2008-2009:

Hut construction  
Crystals production  
Utilities

2010=2011:

Clean room  
External Shielding  
Cryogenics  
CUORE-0

2012:

Internal Shielding  
Detector assembly  
Faraday Cage  
Front-end & DAQ

2013:

Data taking

# CUORE-0

CUORE-0 = first CUORE tower to be installed in the CUORICINO dilution refrigerator (hall A @ LNGS)

## Motivations

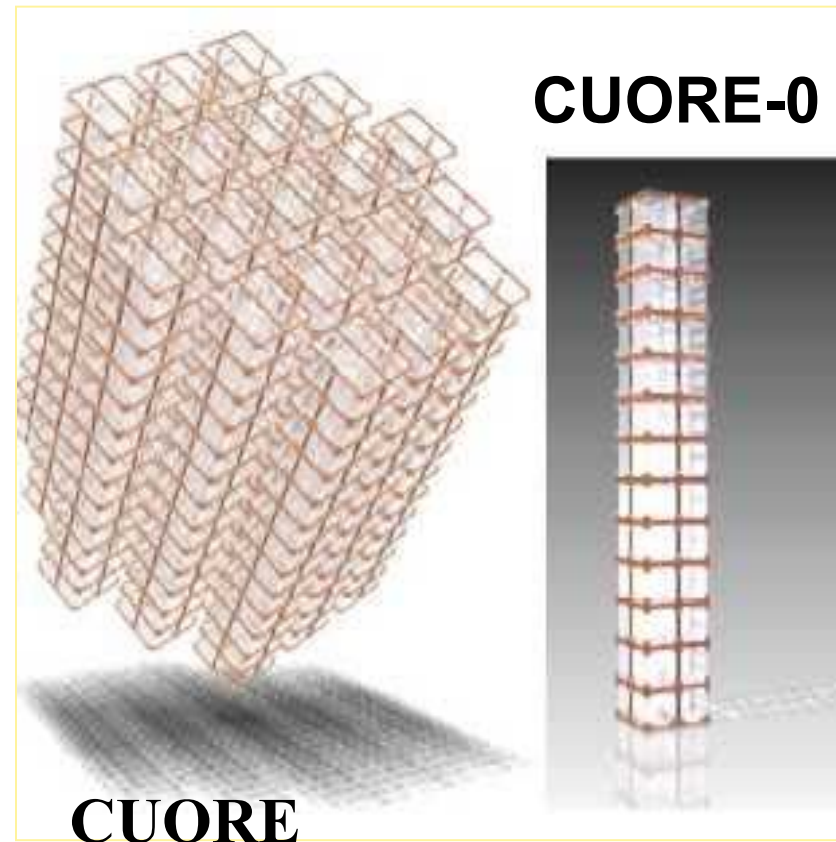
High statistics test of the many improvements/changes developed for the CUORE assembly procedure:

- gluing
- holder
- zero-contact approach
- Wires
- ...

CUORE demonstrator: expected background in the DBD and alpha energy regions reduced by a factor 3 with respect to CUORICINO

**0.07 counts/keV/kg/y**

Powerful experiment: it will overtake soon CUORICINO sensitivity





# CUORE 5 y sensitivity

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- “Realistic”:

$$B = 0.01 \text{ /keV}\cdot\text{kg}\cdot\text{y}; \Delta E = 5 \text{ keV}$$

$$T_{1/2} > 2.1 \cdot 10^{26} \text{ y}, \langle m \rangle < 0.04\text{-}0.09 \text{ eV}$$

- “Optimistic”:

$$B = 0.001 \text{ /keV}\cdot\text{kg}\cdot\text{y}; \Delta E = 5 \text{ keV}$$

$$T_{1/2} > 6.5 \cdot 10^{26} \text{ y}, \langle m \rangle < 0.02\text{-}0.05 \text{ eV}$$

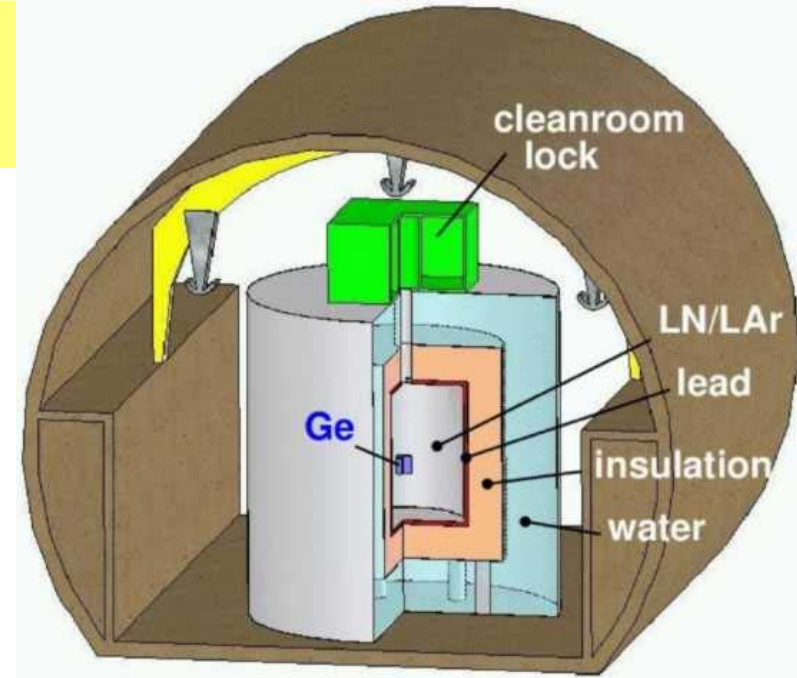
# GERDA

Germany, Italy, Belgium, Russia

Goal: analyse HM evidence in a short time using existing  $^{76}\text{Ge}$  enriched detectors (HM, Igex)

Concept: naked Ge crystals in LAr

- 1.5 m (LAr) + 10 cm Pb + 2 m water
- 2-3 orders of magnitude better bkg than present Status-of-the-Art
- active shielding with LAr scintillation



## 3 phases experiment

Phase I: operate refurbished HM & IGEX enriched detectors (~18 kg)

- Underground commissioning
- Background: 0.01 counts/ keV kg y
- Scrutinize  $^{76}\text{Ge}$  claim with the same nuclide (5s exclusion/confirmation)
- Half life sensitivity:  $3 \times 10^{25}$  y
- Start data taking: 2011

Phase II: additional ~20 kg  $^{76}\text{Ge}$  diodes (segmented detectors)

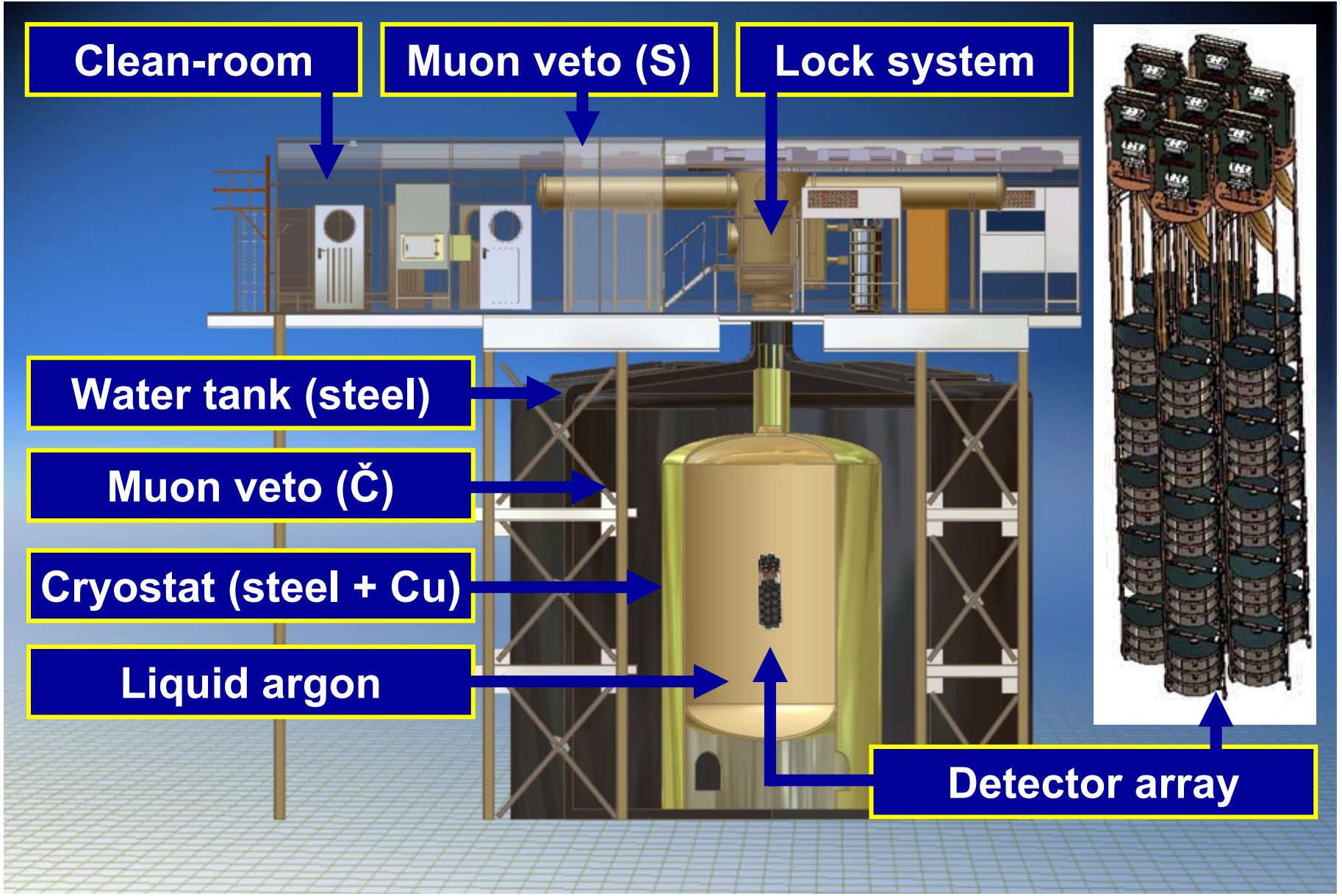
- Background: 0.001 counts / keV kg y
- Sensitivity after 100 kg y (~3 years):  $2 \times 10^{26}$  y ( $\langle m_\nu \rangle < 70 - 200$  meV)

Phase III: depending on physics results of Phase I/II

~ 1 ton experiment in world wide collaboration with MAJORANA

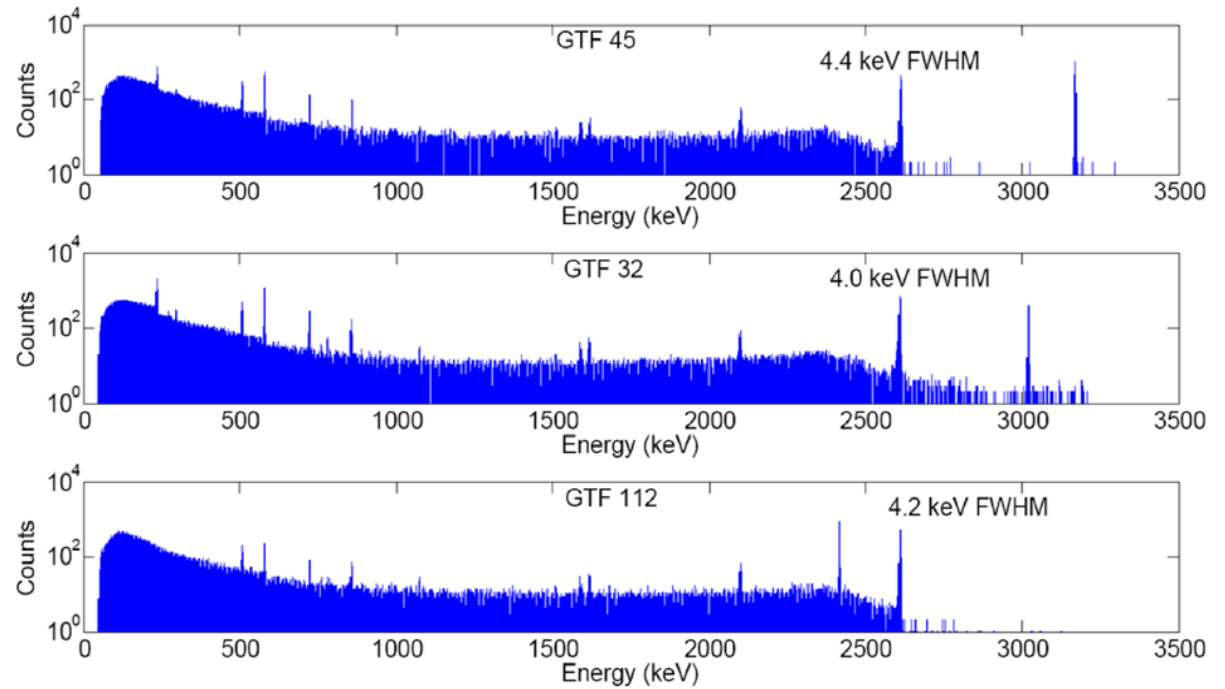
$\langle m_\nu \rangle < 10 - 40$  meV

# GERDA: Technical realization

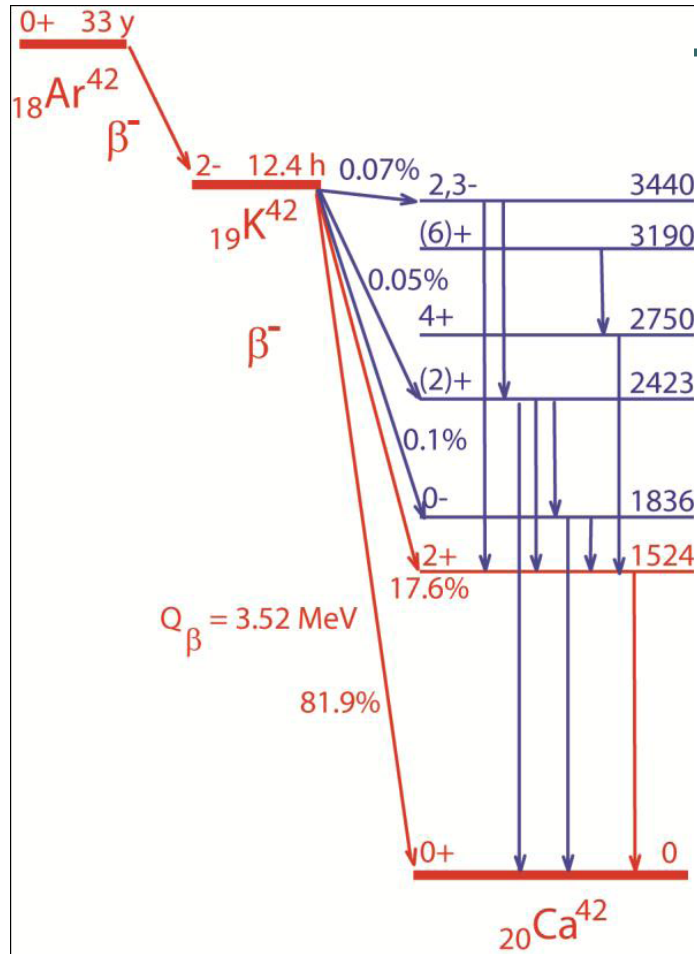


# Operation of the 3 <sup>nat</sup>Ge detectors

## Calibration by <sup>232</sup>Th source



# $^{42}\text{Ar}$ background problem



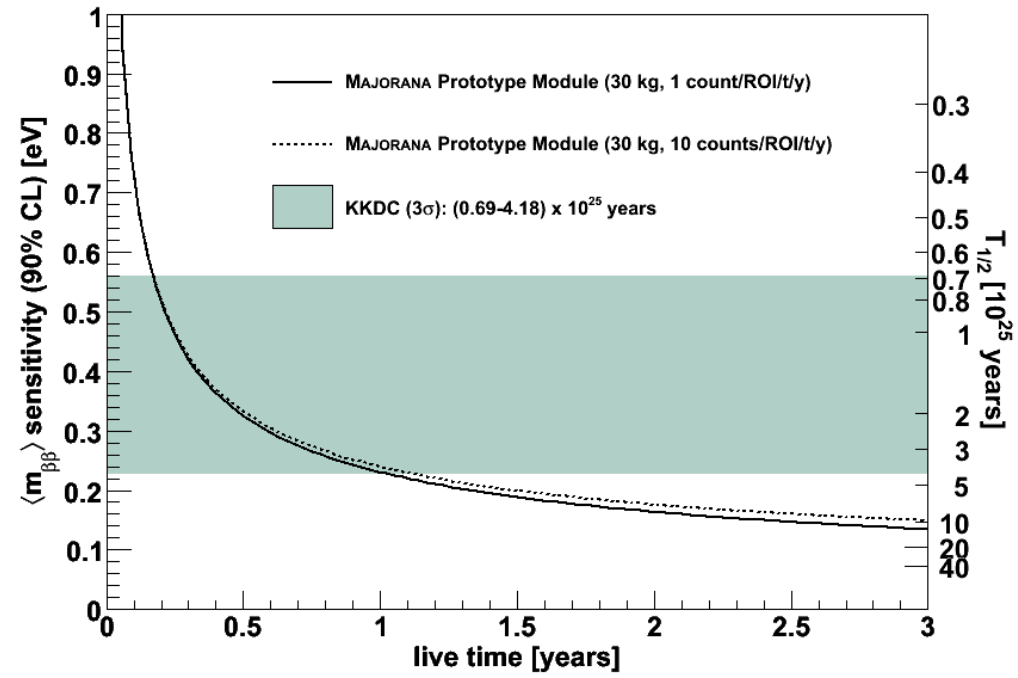
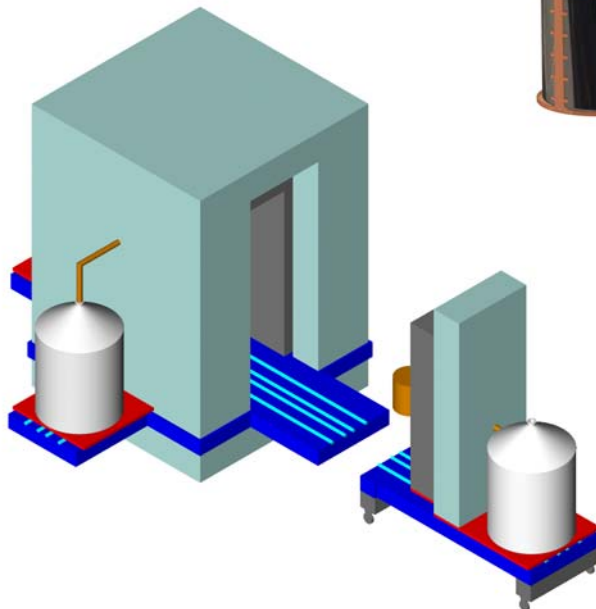
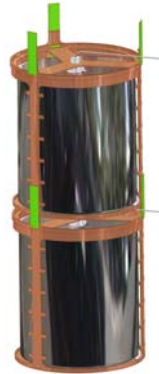
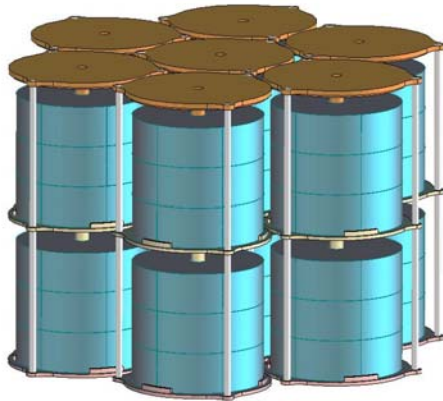
A.S. Barabash et al.,  
 NIM A 416 (1998) 179  
 $^{42}\text{Ar}/^{40}\text{Ar} < 6 \cdot 10^{-21}$  (90% CL)

More than 10 times higher  
 activity in GERDA???



**No contradiction.** GERDA  
 measure not  $^{42}\text{Ar}$ , but local  
 activity of  $^{42}\text{K}$ .  $^{42}\text{K}$  is created as  
 ions and concentrated around Ge  
 detectors, wires and so on  
 because of electric field.

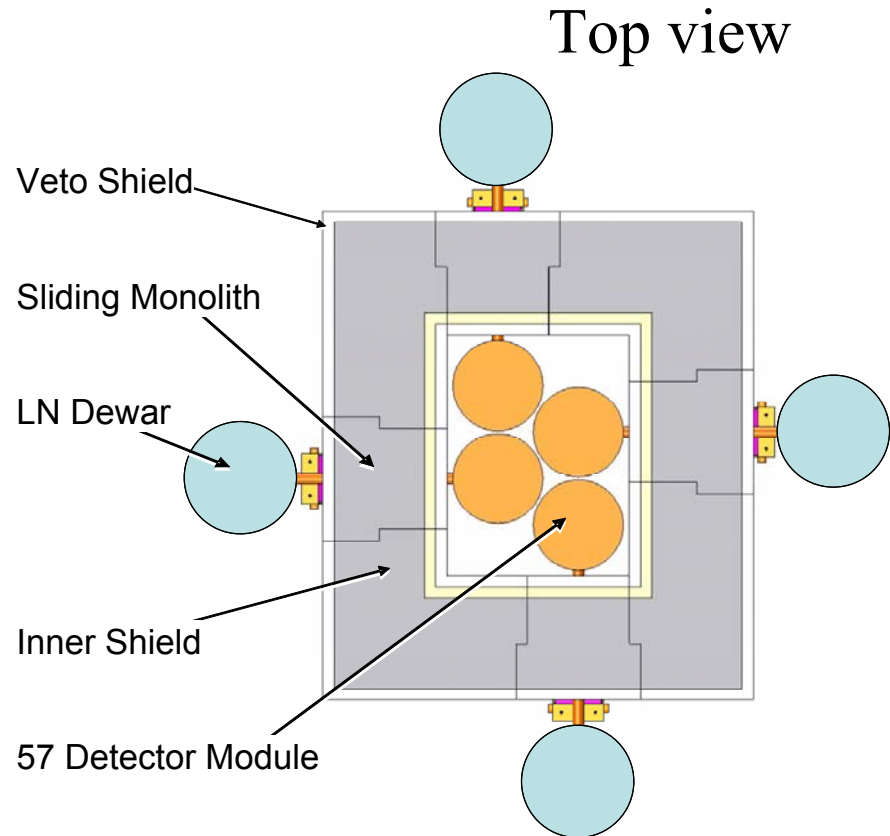
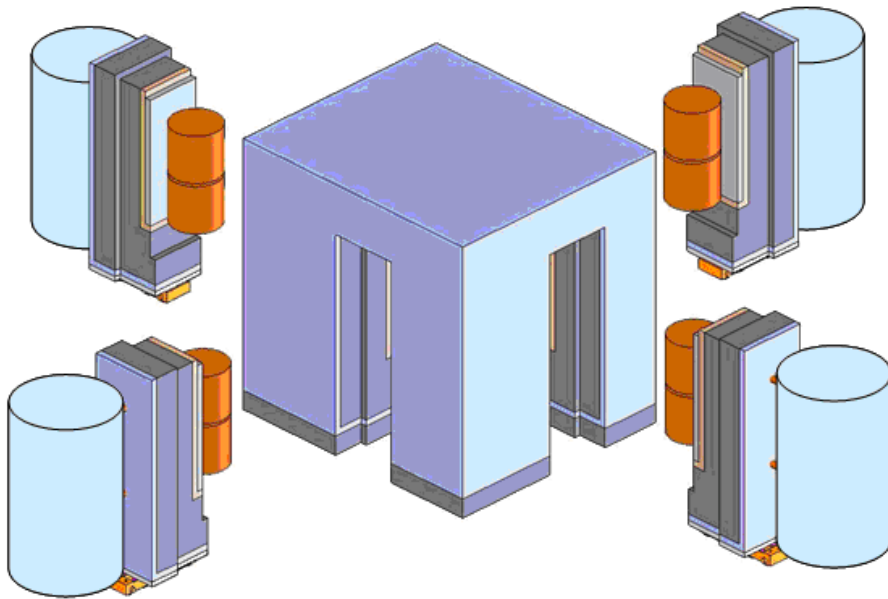
# MAJORANA Project



# The Majorana Shield - Conceptual Design



- Deep underground: >5000'
- Allows modular deployment, early operation
- Contains up to eight 57-crystal modules
- 40 cm bulk Pb, 10 cm ultra-low background shield
- Active  $4\pi$  veto detector



# The MAJORANA DEMONSTRATOR Module



$^{76}\text{Ge}$  offers an excellent combination of capabilities & sensitivities.

(Excellent energy resolution, intrinsically clean detectors, commercial technologies, best  $0\nu\beta\beta$  sensitivity to date)

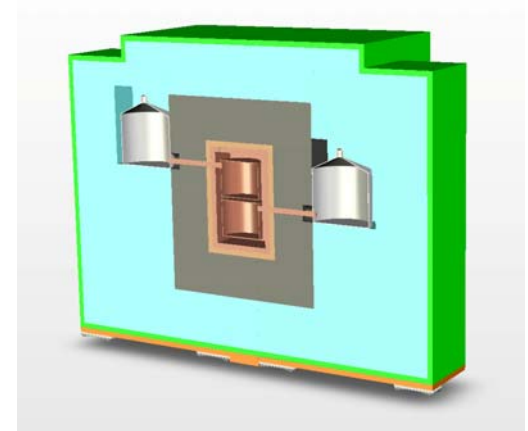
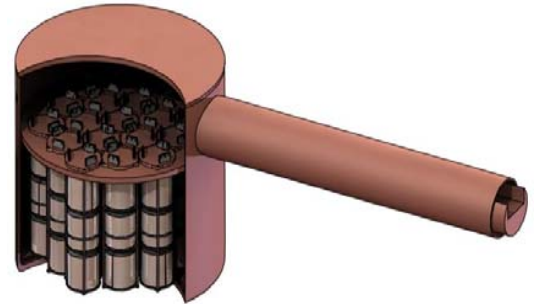
- **60-kg of Ge detectors**

- 30-kg of 86% enriched  $^{76}\text{Ge}$  crystals required for science goal.
- 60-kg required for sensitivity to background goal.
- Examine detector technology options  
p- and n-type, segmentation, point-contact.

- **Low-background Cryostats & Shield**

- ultra-clean, electroformed Cu
- Initial module will have 3 cryostats
- naturally scalable
- Compact low-background passive Cu and Pb shield with active muon veto

- **Located underground 4850' level at SUSEL/DUSEL.**



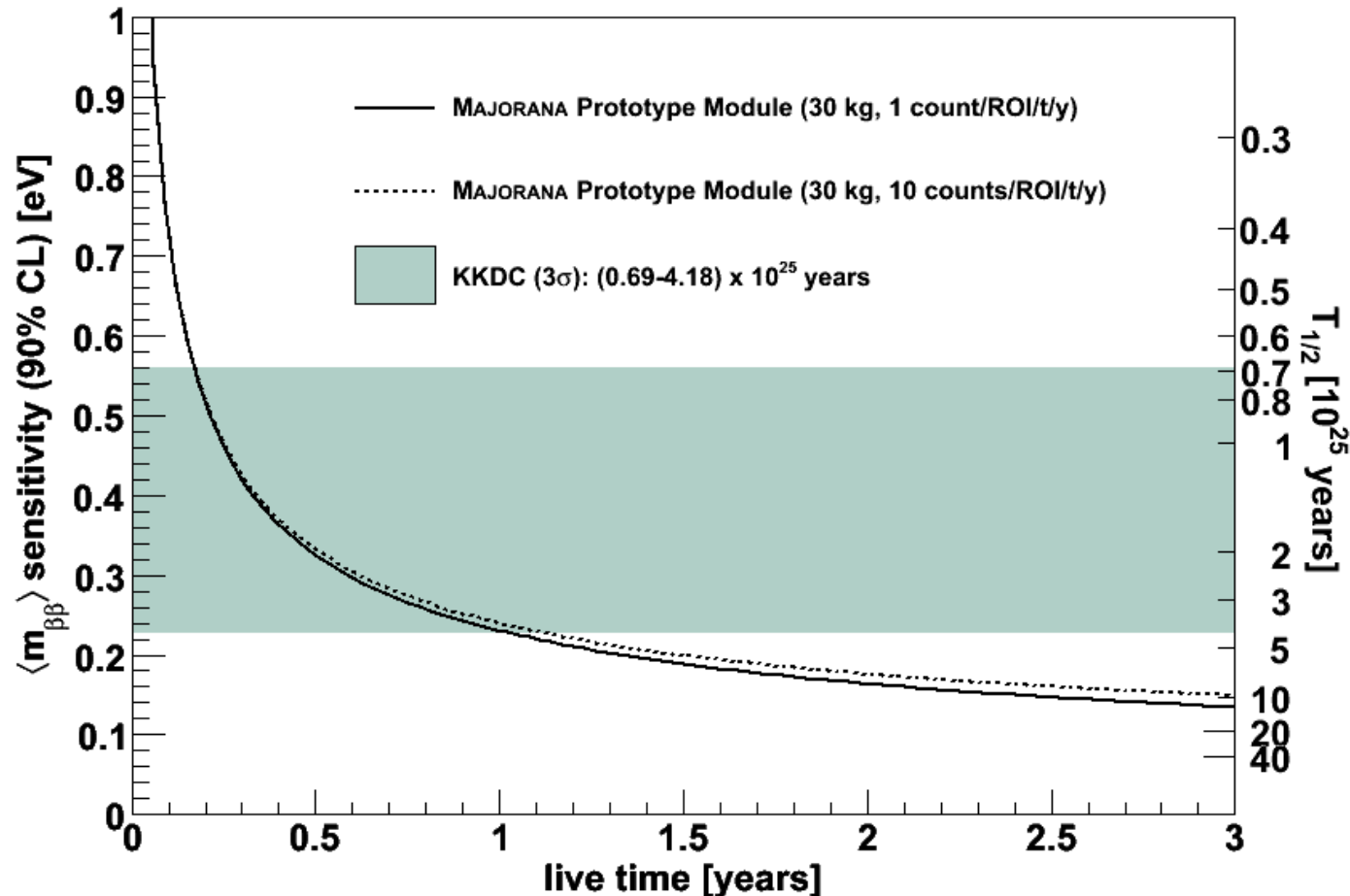


# MAJORANA DEMONSTRATOR Module Sensitivity



- **Expected Sensitivity to  $0\nu\beta\beta$**   
(30 kg enriched material, running 3 years, or 0.09 t-y of  $^{76}\text{Ge}$  exposure)

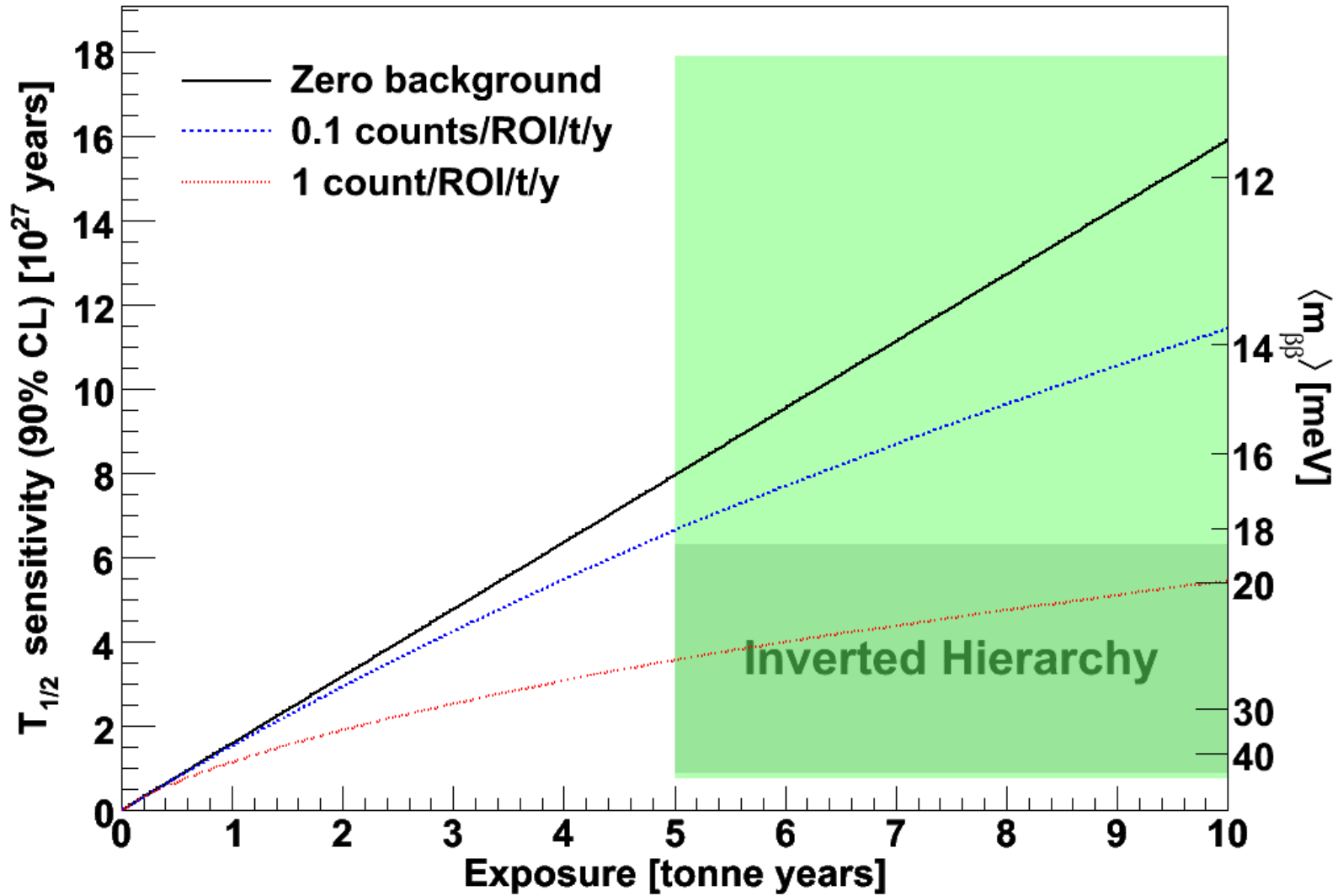
$T_{1/2} \geq 10^{26}$  y (90% CL). Sensitivity to  $\langle m_{\nu} \rangle < 140$  meV (90% CL) [Rod05,err.]



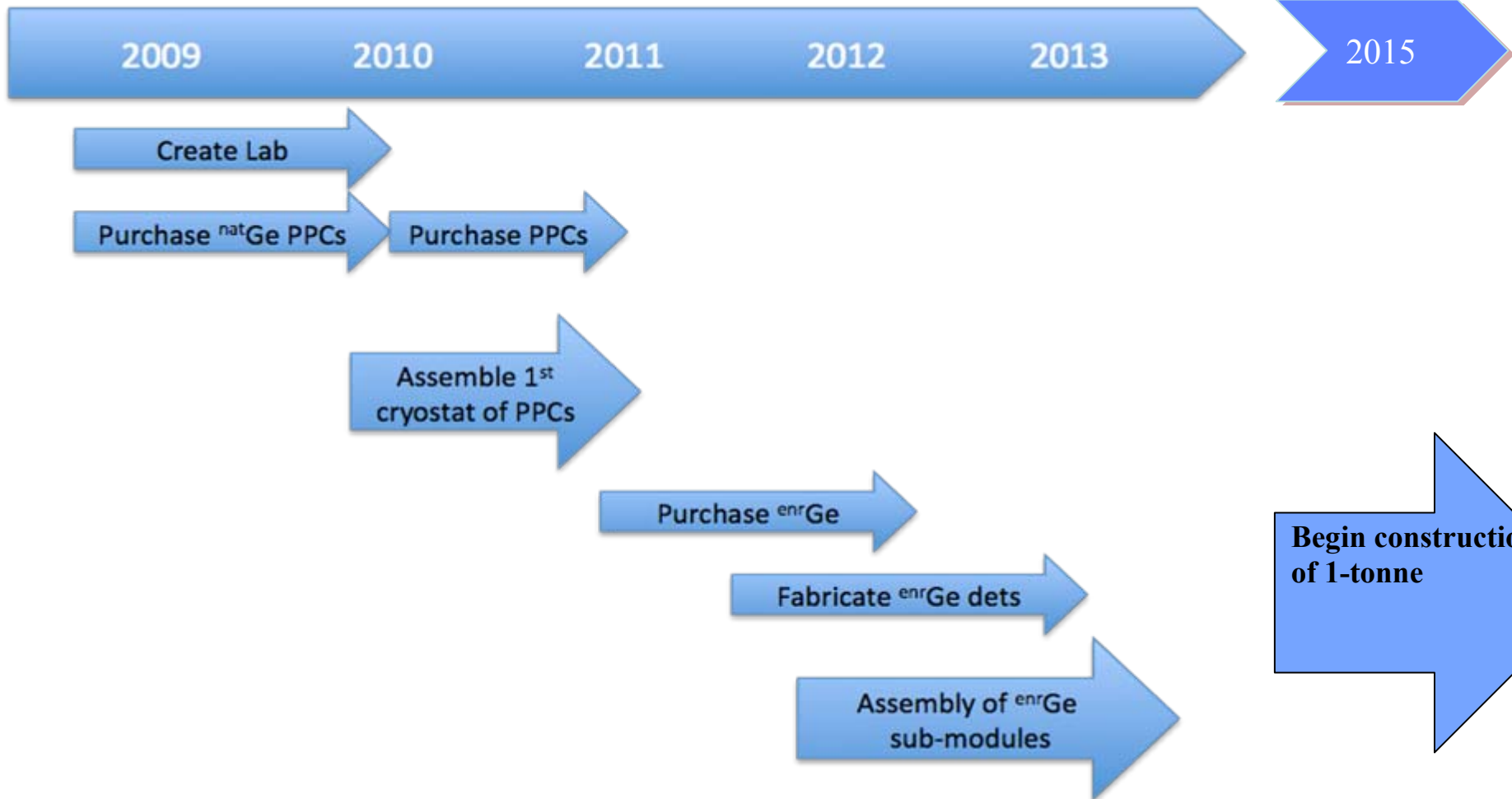
# 1-tonne Ge - Projected Sensitivity vs. Background



$$T_{1/2}^{0\nu} = \ln(2)N\epsilon t/UL(B)$$



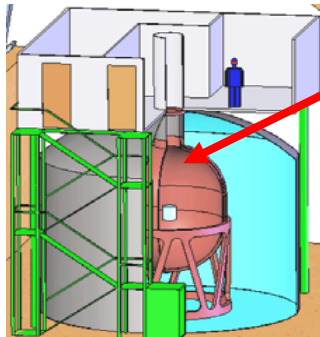
# MAJORANA DEMONSTRATOR SCHEDULE



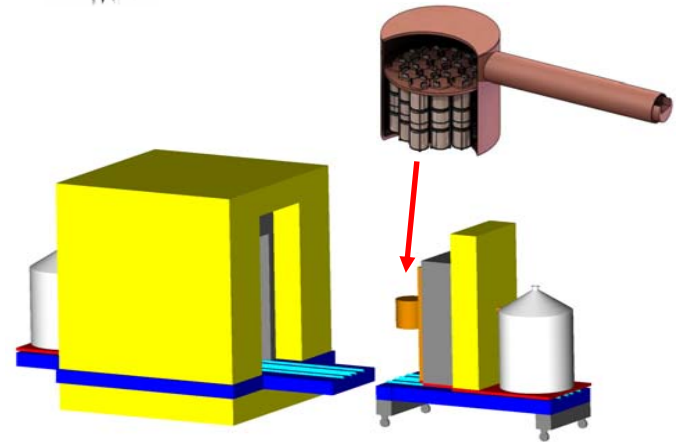
# GERDA - Majorana



**GERDA**



**Majorana**



- 'Bare'  $^{76}\text{Ge}$  array in liquid argon
- Shield: high-purity liquid Argon /  $\text{H}_2\text{O}$
- Phase I (~2011): ~18 kg (HdM/IGEX diodes)
- Phase II (~2012): add ~20 kg new detectors  
Total ~40 kg

- Modules of  $^{76}\text{Ge}$  housed in high-purity electroformed copper cryostat
- Shield: electroformed copper / lead
- Initial phase: R&D prototype module  
Total 60 kg

## **Joint Cooperative Agreement:**

- Open exchange of knowledge & technologies (e.g. MaGe, R&D)
- Intention to merge for 1 ton exp. Select best techniques developed and tested in GERDA and Majorana

# EXO (Enriched Xenon Observatory)

USA-RUSSIA-CANADA

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- $^{136}\text{Xe} \rightarrow ^{136}\text{Ba}^{++} + 2e^{-}$  ( $E_{2\beta} = 2.47 \text{ MeV}$ )
- **Main idea is:** to detect all products of the reaction with good enough energy and space resolution (M.Moe PRC 44(1991)931)



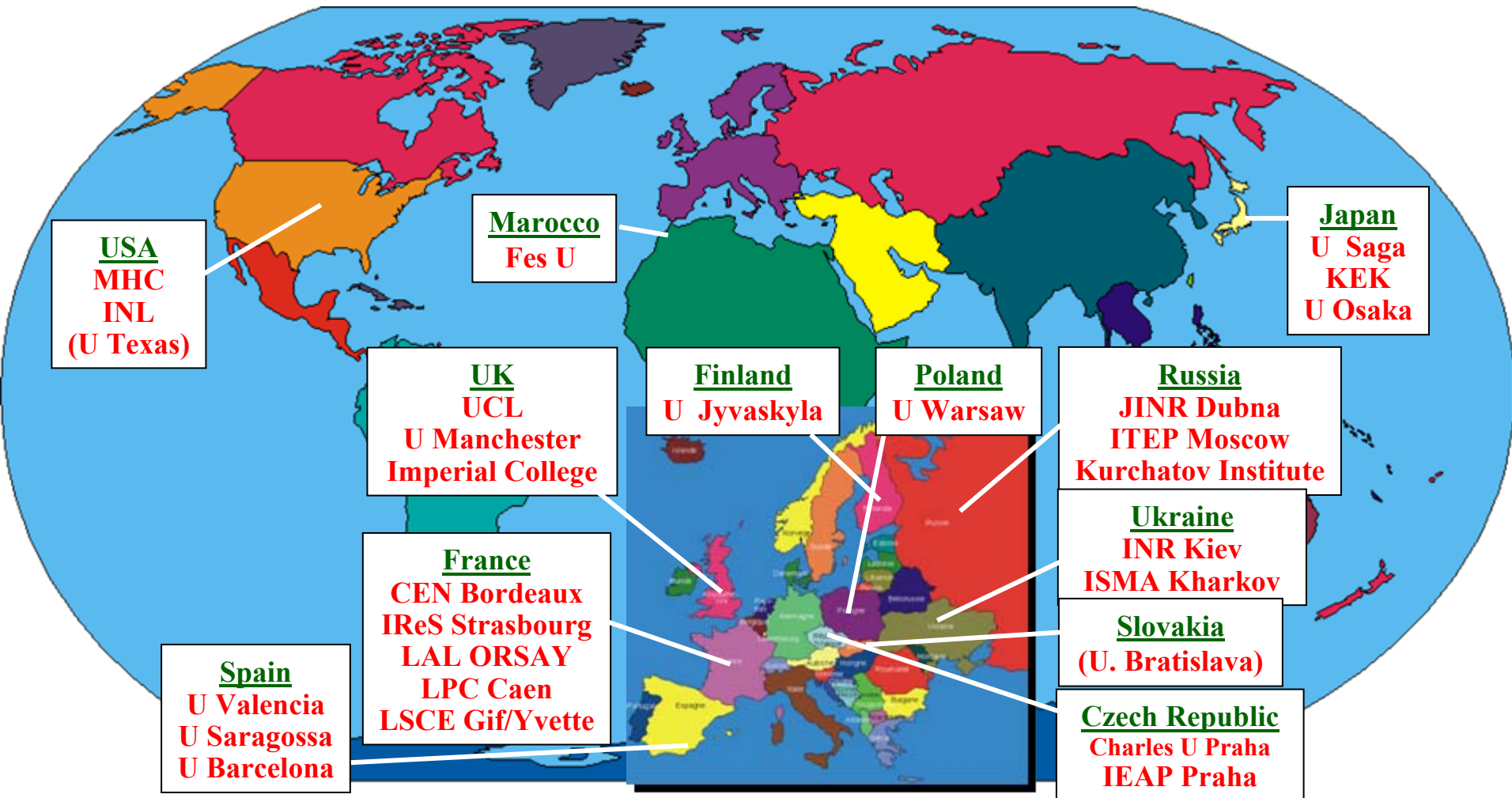
## EXO-200 (without Ba<sup>+</sup> tagging)

---

- **200 kg** of <sup>136</sup>Xe (80% enrichment) – **exist!**
- Location: WIPP (USA)
- $\Delta E/E(\text{FWHM}) = \mathbf{3.8\%}$  at 2.5 MeV (ionization and scintillation readout)
- Background (5 y) = **40** events
- Sensitivity (5 y):  **$6.4 \cdot 10^{25}$  y** (  **$\langle m_\nu \rangle \sim 0.1-0.2$  eV** )
- Start of measurements: in  **$\sim 2011$**

# SuperNEMO Collaboration

~ 90 physicists, 12 countries, 27 laboratories



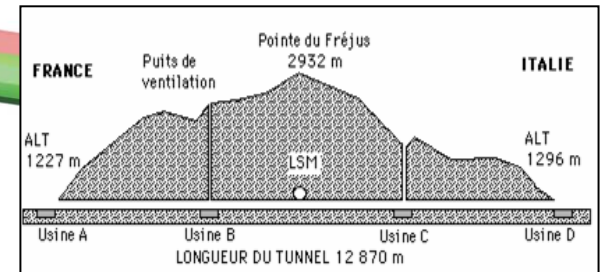
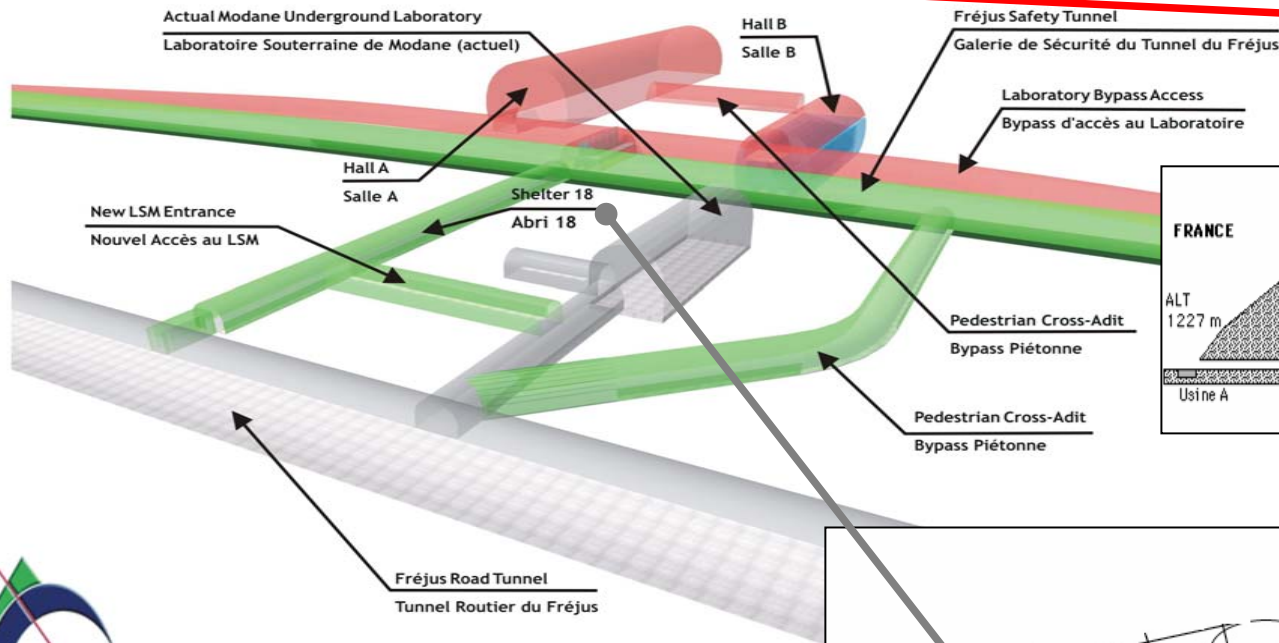


# Possible location : LSM

MODANE UNDERGROUND LABORATORY 60'000 m<sup>3</sup> EXTENSION

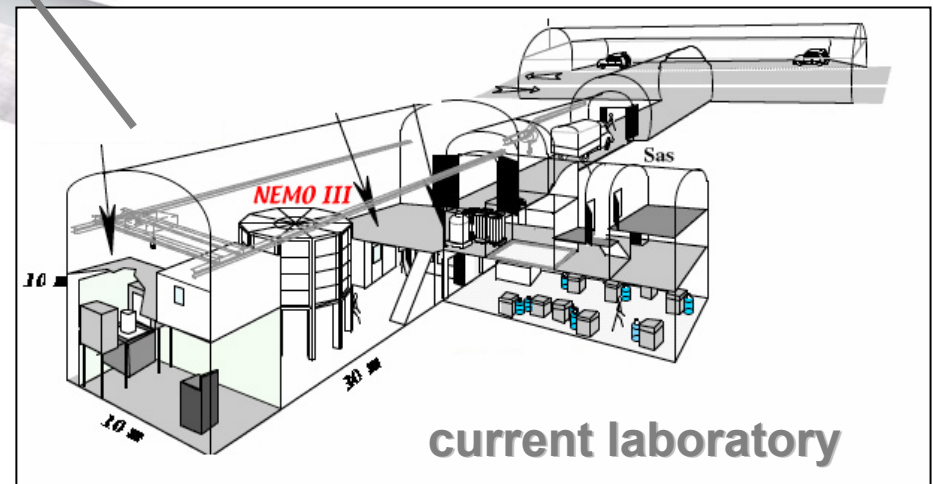
LABORATOIRE SOUTERRAINE DE MODANE AGRANDISSEMENT 60'000 m<sup>3</sup>

**A 60 000 m<sup>3</sup>  
extension**

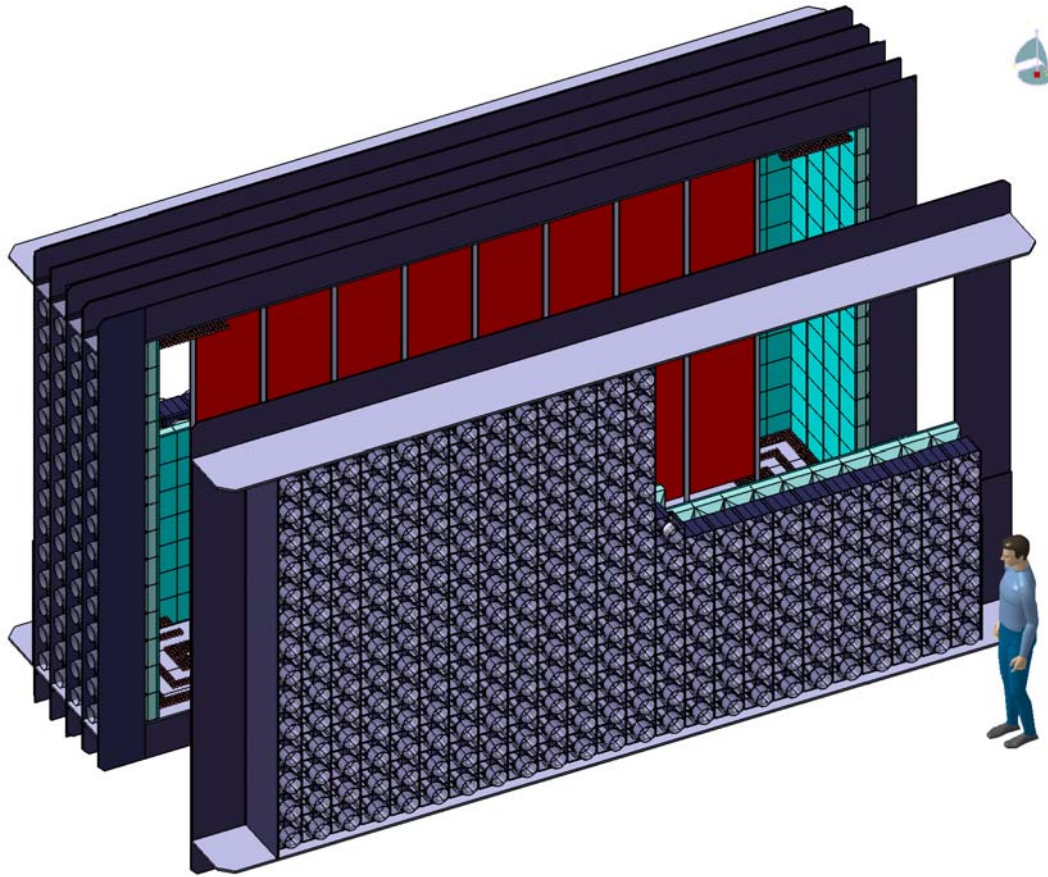


Demonstrator to be placed in NEMO3's current location

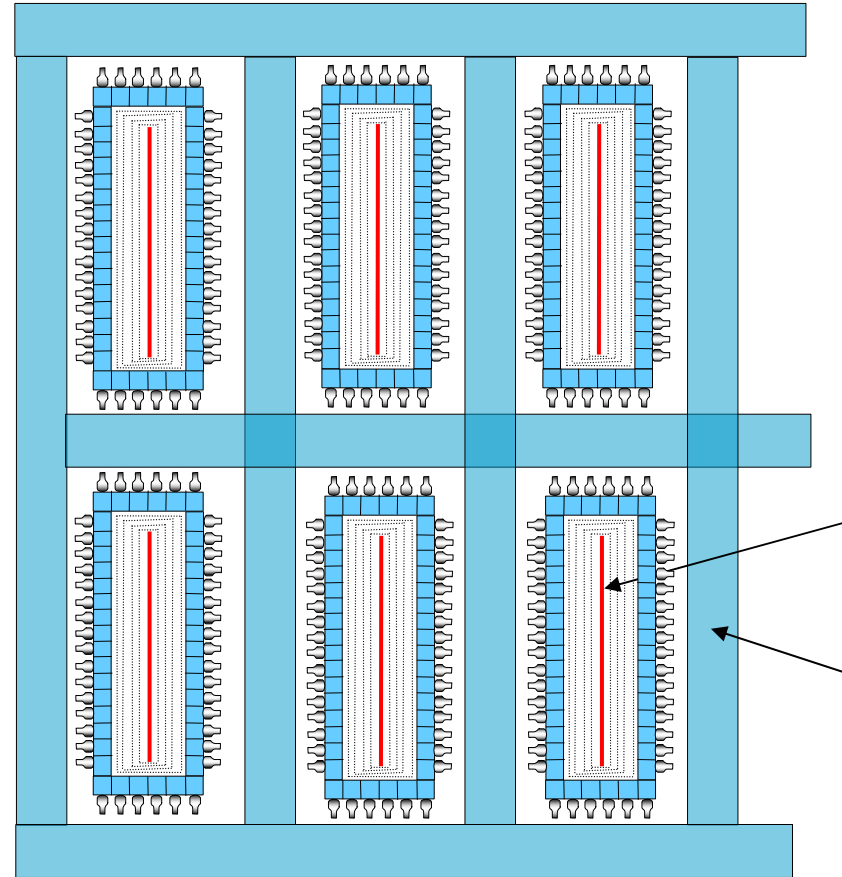
Full SuperNEMO detector in a new cavity (extension finished on 2012)



# Very preliminary design



Single sub-module  
with ~7 kg of isotope



~20 sub-modules for 100+ kg of isotope  
surrounded by shielding

# From NEMO-3 to SuperNEMO

## NEMO-3

## SuperNEMO

$^{100}\text{Mo}$	Choice of isotope	$^{82}\text{Se}$ or $^{150}\text{Nd}$
7 kg	Isotope mass M	100-200 kg
8% @3MeV	Energy resolution FWHM (calorimeter)	4% @ 3MeV
20 %	Efficiency $\mathcal{E}(\beta\beta 0\nu)$	~ 30 %
$^{208}\text{Tl} < 20 \mu\text{Bq/kg}$ $^{214}\text{Bi} < 300 \mu\text{Bq/kg}$	Internal radiopurity <i><math>^{208}\text{Tl}</math> and <math>^{214}\text{Bi}</math></i> in the $\beta\beta$ foils	$^{208}\text{Tl} < 2 \mu\text{Bq/kg}$ (If $^{82}\text{Se}$ : $^{214}\text{Bi} < 10 \mu\text{Bq/kg}$ )
$T_{1/2}(\beta\beta 0\nu) > 2 \cdot 10^{24} \text{ y}$ $\langle m_\nu \rangle < 0.3 - 1.3 \text{ eV}$	SENSITIVITY	$T_{1/2}(\beta\beta 0\nu) > (1-2) \cdot 10^{26} \text{ y}$ $\langle m_\nu \rangle \sim 40-110 \text{ meV}$

Main R&D tasks:

- 1)  $\beta\beta$  source production
- 2) Energy resolution

- 3) Radiopurity
- 4) Tracking

# SuperNEMO Demonstrator (1<sup>st</sup> module)

## MAIN GOALS :

- To demonstrate the feasibility of large scale detector with required performance (efficiency, energy resolution, radiopurity, ...)

- To measure the radon background

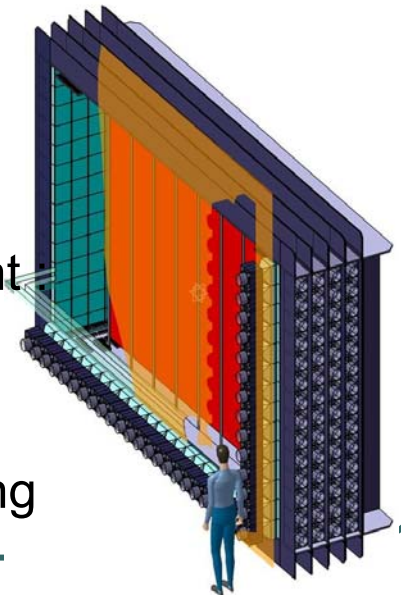
- To finalize detector design

- To produce competitive physics measurement

$$T_{1/2}(\beta\beta 0\nu) > 6.5 \times 10^{24} \text{ years}$$

$$\langle m_\nu \rangle < 210 - 570 \text{ meV}$$

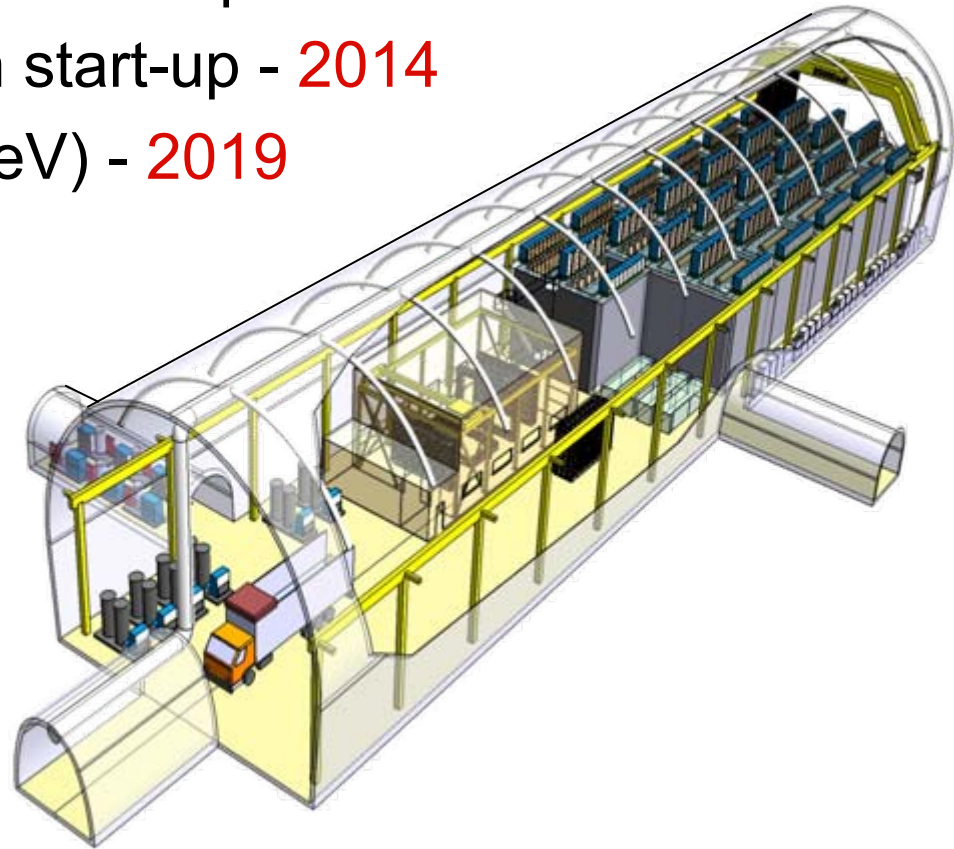
with 7 kg of  $^{82}\text{Se}$  after  $\sim 2$  years of demonstrator data taking



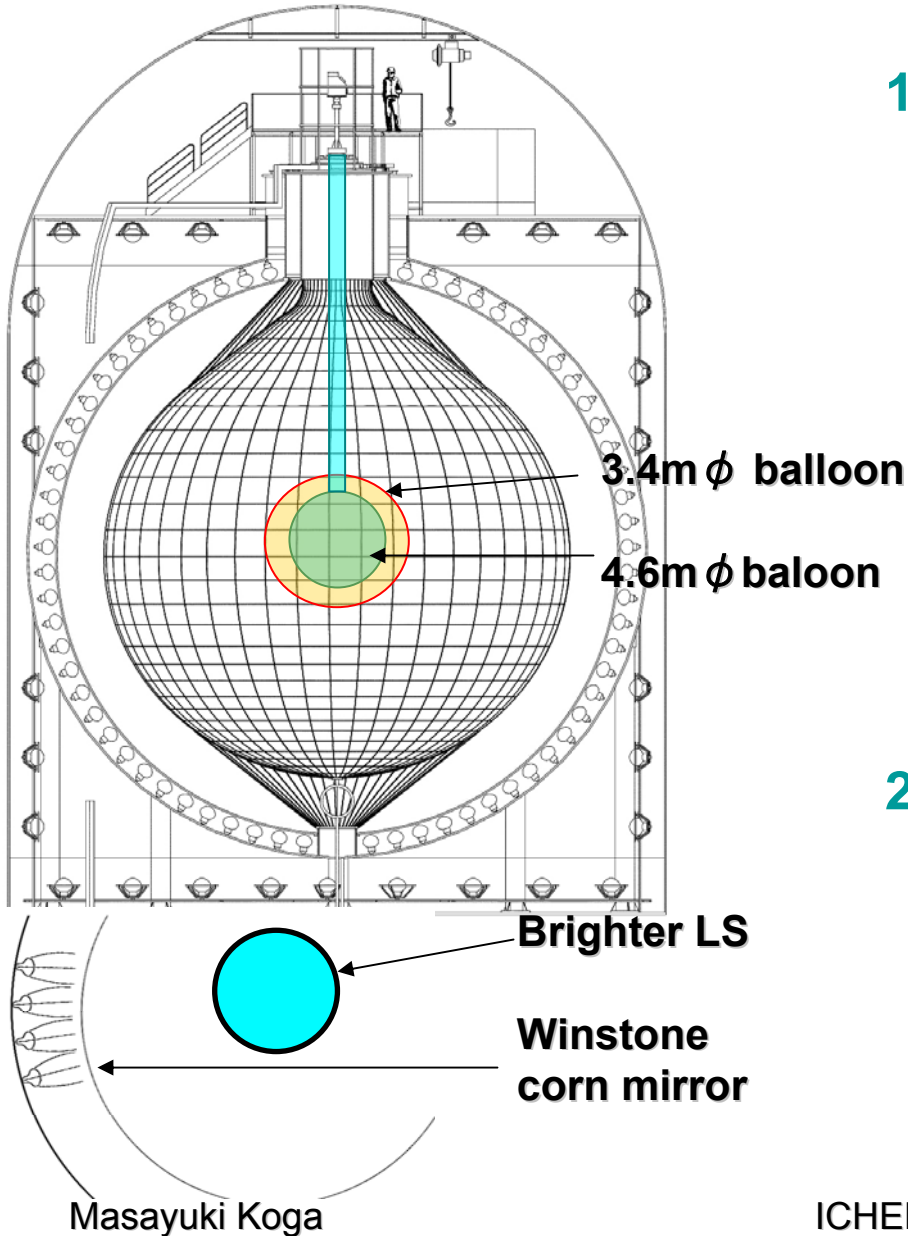
# SuperNEMO schedule highlights

- NEMO-3 decommissioning - early 2011
- Demonstrator construction - 2010-2012
- Demonstrator physics run start-up - 2013
- Full detector construction start-up - 2014
- Target sensitivity ( $\sim 0.05$  eV) - 2019

KK claim to be verified with Demonstrator by 2015



# KamLAND-Zen project



**1st phase enriched Xe 400kg**

R=1.7m balloon

V=20.5m<sup>3</sup>, S=36.3m<sup>2</sup>

LS : C10H22(81.8%)+PC(18%)

+PPO+Xe(~2.5wt%)

$\rho$  LS: 0.78kg/ℓ

high sensitivity with low cost



tank opening (2013 or 2015)

**2nd phase enriched Xe 1000kg**

R=2.3m balloon

V=51.3m<sup>3</sup>, S=66.7m<sup>2</sup>

improvement of energy resolution

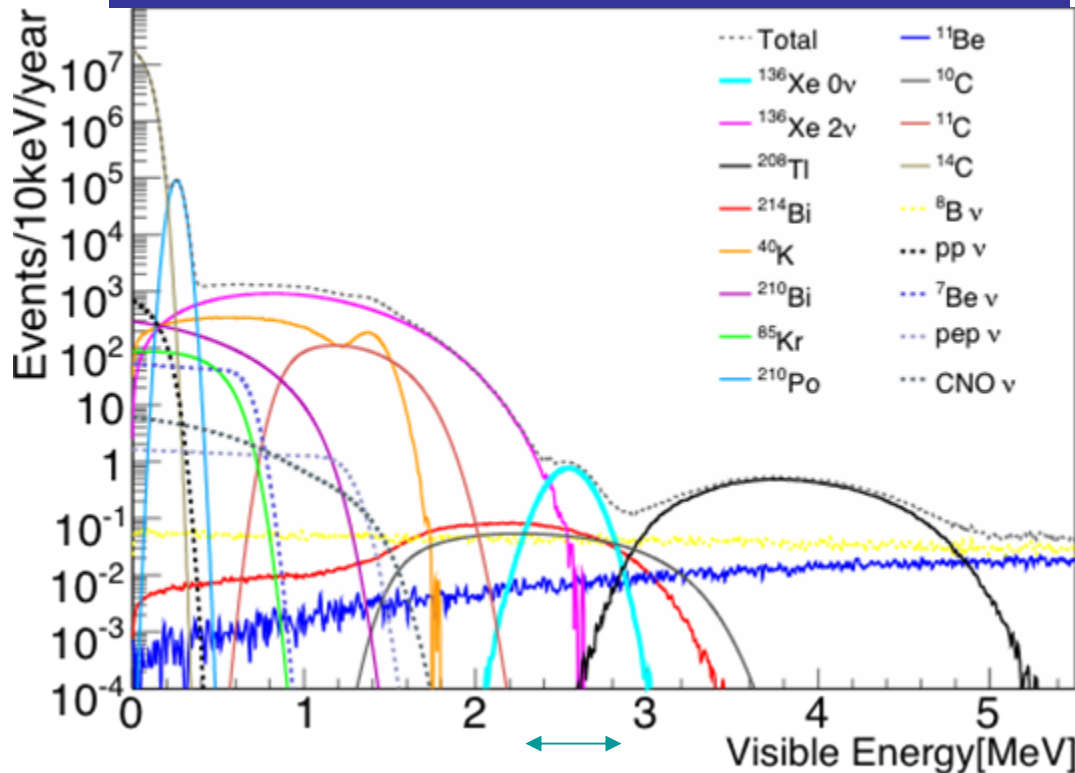
(brighter LS, higher light concentrator )

# Background study using KamLAND MC (GEANT4)

## Major BG

- (1).  $^{136}\text{Xe } 2\nu \beta\beta$
- (2). spallation isotopes :  $^{10}\text{C}$ ,  $^{11}\text{Be}$  => 1/10 using new electronics help
- (3).  $^8\text{B}$  solar neutrinos <4.9 events/d/kton on KamLAND
- (4). from Mini Balloon (MIB) material :  $^{208}\text{Tl}$ ,  $^{214}\text{Bi}$  => vertex cut,

## Simulated Energy Spectrum at KamLAND



## Assumed

- 400kg 90% enriched Xe loaded LS
- MIB contamination ( $^{238}\text{U}$ ,  $^{232}\text{Th}$ , 40K)  
= (10-12, 10-12, 10-11)[g/g]
- neutrino effective mass  $\langle m\nu \rangle$   
= 150meV (the lower limit of the current claimed detection)
- $T_{1/2}(2\nu\beta\beta) > 10^{22}\text{y}$
- $T_{1/2}(0\nu\beta\beta) > 1.14 \times 10^{24}\text{y}$
- $^{10}\text{C}$  90% tag,  $^{214}\text{Bi}$  66% tag

## Summary of BG and signal in signal region

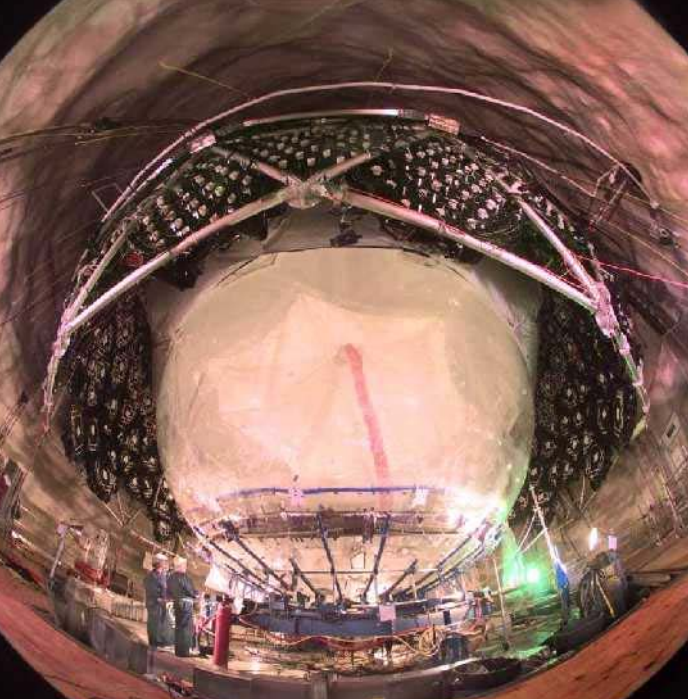
$^{136}\text{Xe } 2\nu$	$^{208}\text{Tl}$	$^{214}\text{Bi}$	$^{10}\text{C}$	$^{11}\text{Be}$	$^8\text{B}$	Total	$^{136}\text{Xe } 0\nu$
2.08	$1.86 \times 10^{-2}$	2.40	3.09	0.26	1.52	9.35	18.08
$\pm 0.15$	$\pm 0.13 \times 10^{-2}$	$\pm 0.01$	$\pm 0.01$	$\pm 0.01$	$\pm 0.03$	$\pm 0.23$	$\pm 0.02$

[events/year]

# summary

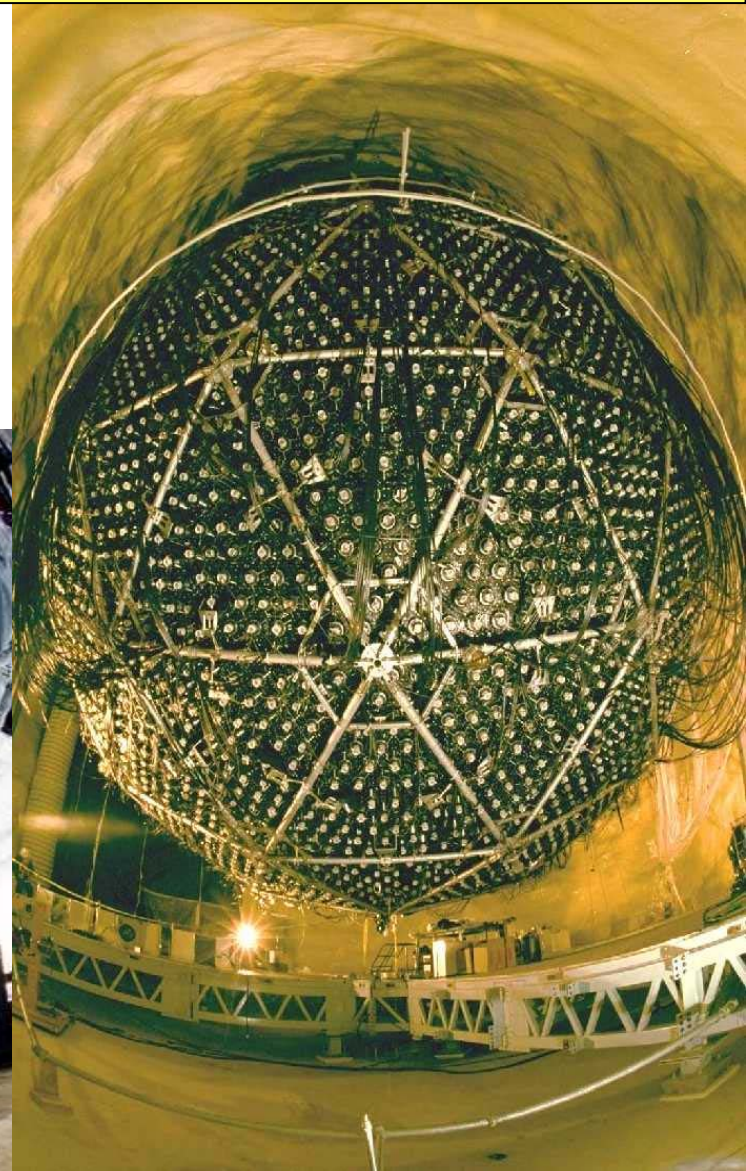
- KamLAND is running for reactor, Geo,  $7\text{Be}$  solar (to 2011)
- KamLAND have ability to do  $0\nu\beta\beta$  experiment
- KamLAND-Zen project will start using 400kg 90% enriched Xe from **May 2011**
- Target sensitivity on **400kg Phase ~60meV @2years**
- Planning **Xe1000** phase (from **2013** or **2015**: depend on funding)





**SNO: One million pieces transported down in the 9 ft x 12 ft x 9 ft mine cage and re-assembled under ultra-clean conditions. Every worker takes a shower and wears clean, lint-free clothing.**

**Over 70,000 Showers to date and counting**



# SNO+

## SNO+: SNO filled with liquid scintillator

A liquid scintillator detector has poor energy resolution

Huge quantities of isotope (high statistics) and low backgrounds however help compensate

- source in–source out capability
- large, homogeneous liquid detector leads to well-defined background model
- possibly source in–source out capability
- using the technique that was developed originally for LENS and now also used for Gd-loaded scintillator
- SNO+ collaboration managed to load Nd into pseudocumene and in linear alkylbenzene (>1% concentration)
- with 1% Nd loading (natural Nd) a very good neutrinoless double beta decay sensitivity is predicted, but...

### Nd loaded scintillator:

1% loading (Natural Nd) large light absorption by Nd

**$47 \pm 6$  pe/MeV (Monte Carlo)**

0.1% loading (Isotopically enriched to 56% Nd) acceptable

**$400 \pm 21$  pe/MeV (Monte Carlo)**

# SNO+ (2)

- Using existing SNO infrastructure
- Well understood detector

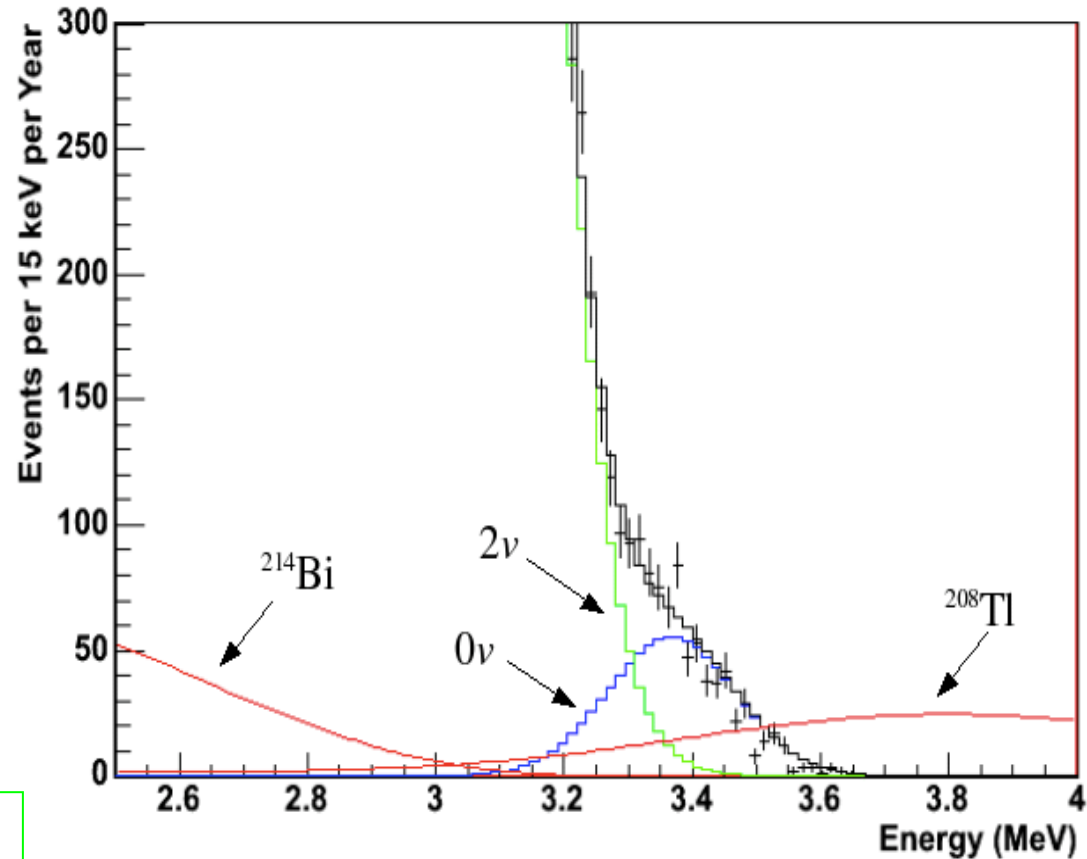
1057 events per year with 500 kg  $^{150}\text{Nd}$ -loaded liquid scintillator in SNO+.

Simulation assuming light output and background similar to Kamland.

## Sensitivity Limits (3 yrs):

- Natural Nd (56 kg isotope):  
 $m_\nu \sim 0.1\text{-}0.3 \text{ eV}$
- 500 kg enriched  $^{150}\text{Nd}$   
 $m_\nu \sim 0.04\text{-}0.12 \text{ eV}$

The Simulated Spectrum of Double Beta Decay Events



Funded by NSERC for final design/engineering and initial construction 2008-2010  
End of 2010 → ready for scintillator filling

## IV. Заключение

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1. Большие успехи достигнуты в изучении  $2\nu$ -распада.
2. Современное консервативное ограничение на  $\langle m_\nu \rangle$  из экспериментов по  $2\beta(0\nu)$ -распаду составляет  $0.75 \text{ eV}$ .
3. Существует указание на «наблюдение»  $2\beta(0\nu)$ -распада в  $^{76}\text{Ge}$  ( $\langle m_\nu \rangle \approx 0.3-0.5 \text{ eV}$ ). Но это требует подтверждения («закрытия») в новых экспериментах с  $^{76}\text{Ge}$  (это будет сделано в ближайшие несколько лет).
4. В 2011 году ожидается:
  - начало набора данных на установке **GERDA-I** (18 кг  $^{76}\text{Ge}$ );
  - начало набора данных на установке **EXO** (200 кг  $^{136}\text{Xe}$ );
  - начало набора данных на установке **KamLAND-Xe** (400 кг  $^{136}\text{Xe}$ );
  - начало набора данных на установке **CUORE-0** (11 кг  $^{130}\text{Te}$ ).
5. В экспериментах нового поколения чувствительность к  $\langle m_\nu \rangle$  на уровне  $\sim (0.01-0.1) \text{ эВ}$  будет достигнута в  $\sim 2013-2020$ .